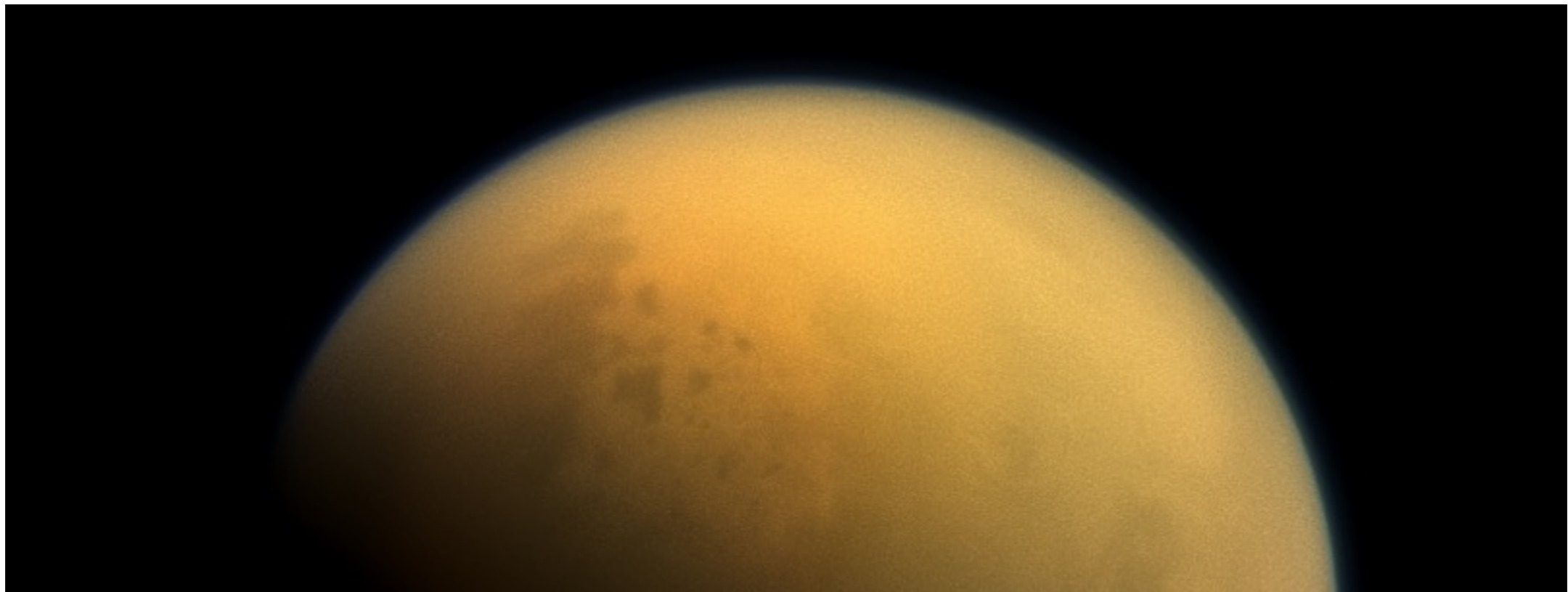


# VUV photolysis of molecules in astrophysical media: Branching ratios of photodissociative channels for databases

Bérenger Gans

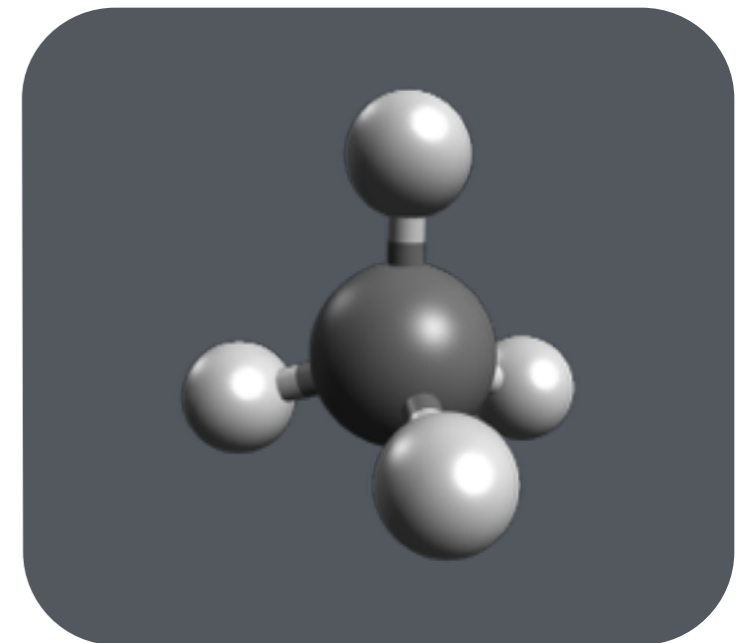
*Institut des Sciences Moléculaires d'Orsay, Université Paris-Sud*



# Outline

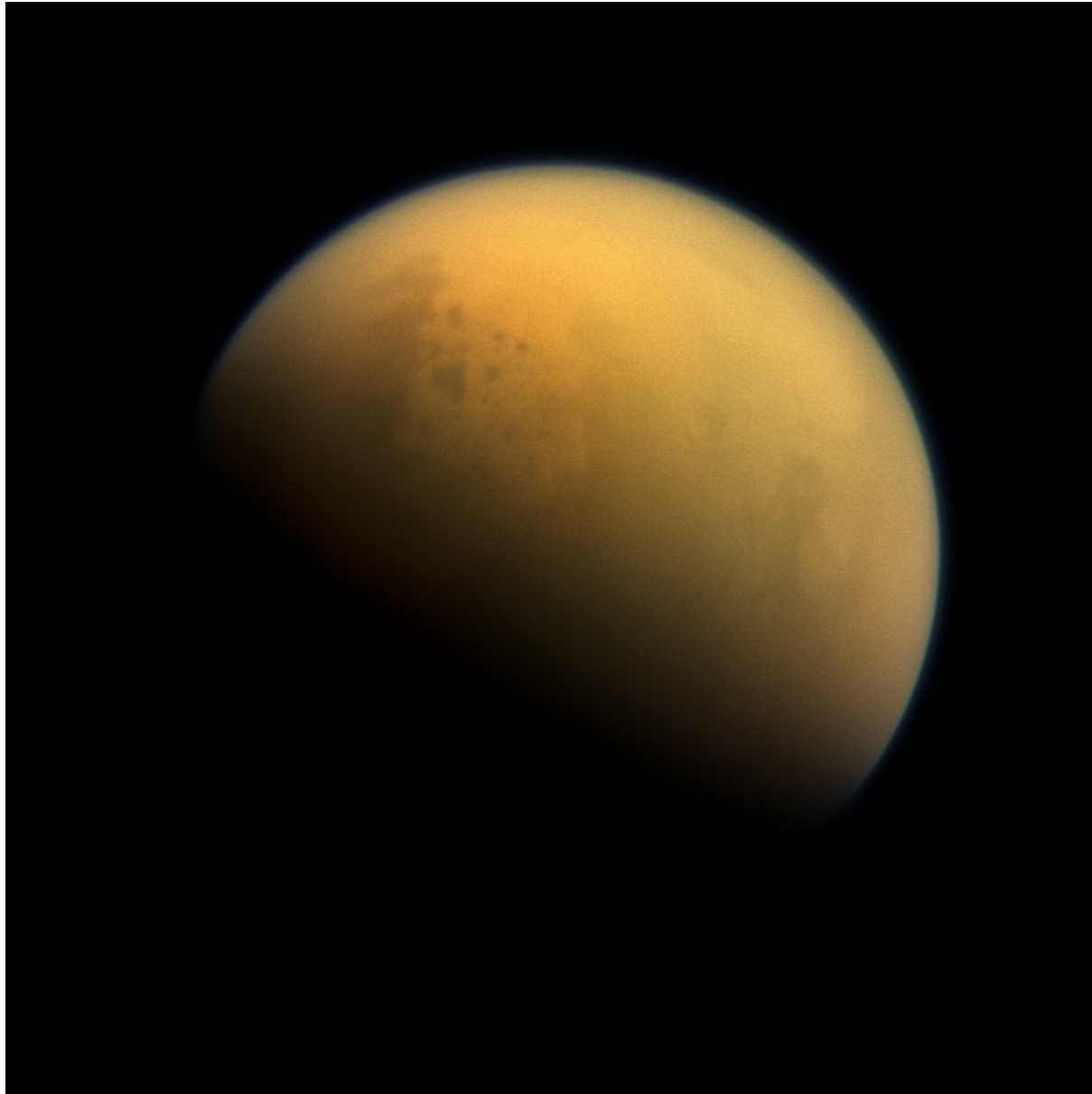
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- Introduction and motivation in the context of Titan
- Example: VUV Photolysis of Methane ( $\text{CH}_4$ )
  - First step: radical photoionization cross section measurement.
  - Second step: Branching ratio measurement
- Conclusions and perspectives



# Introduction

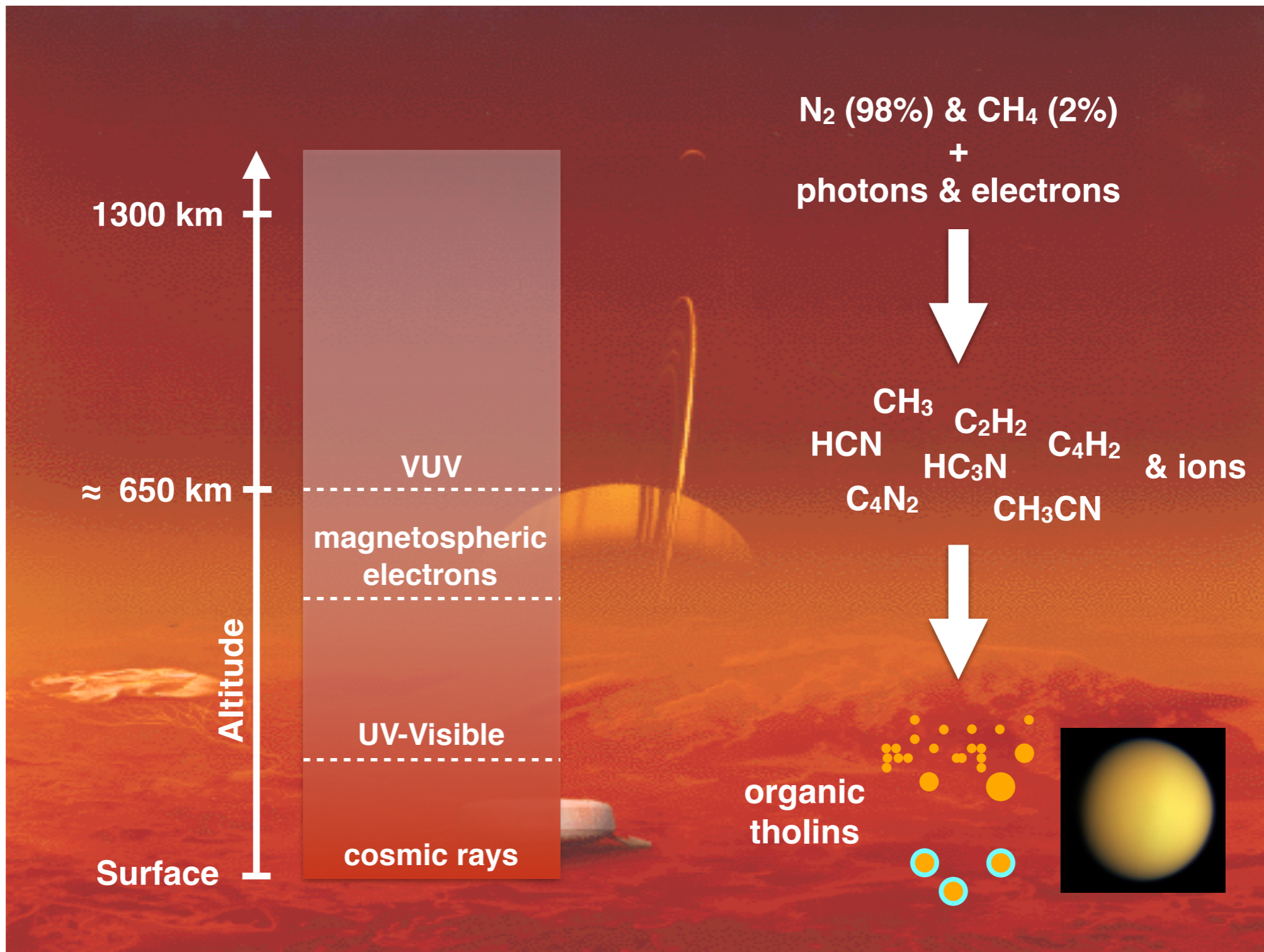
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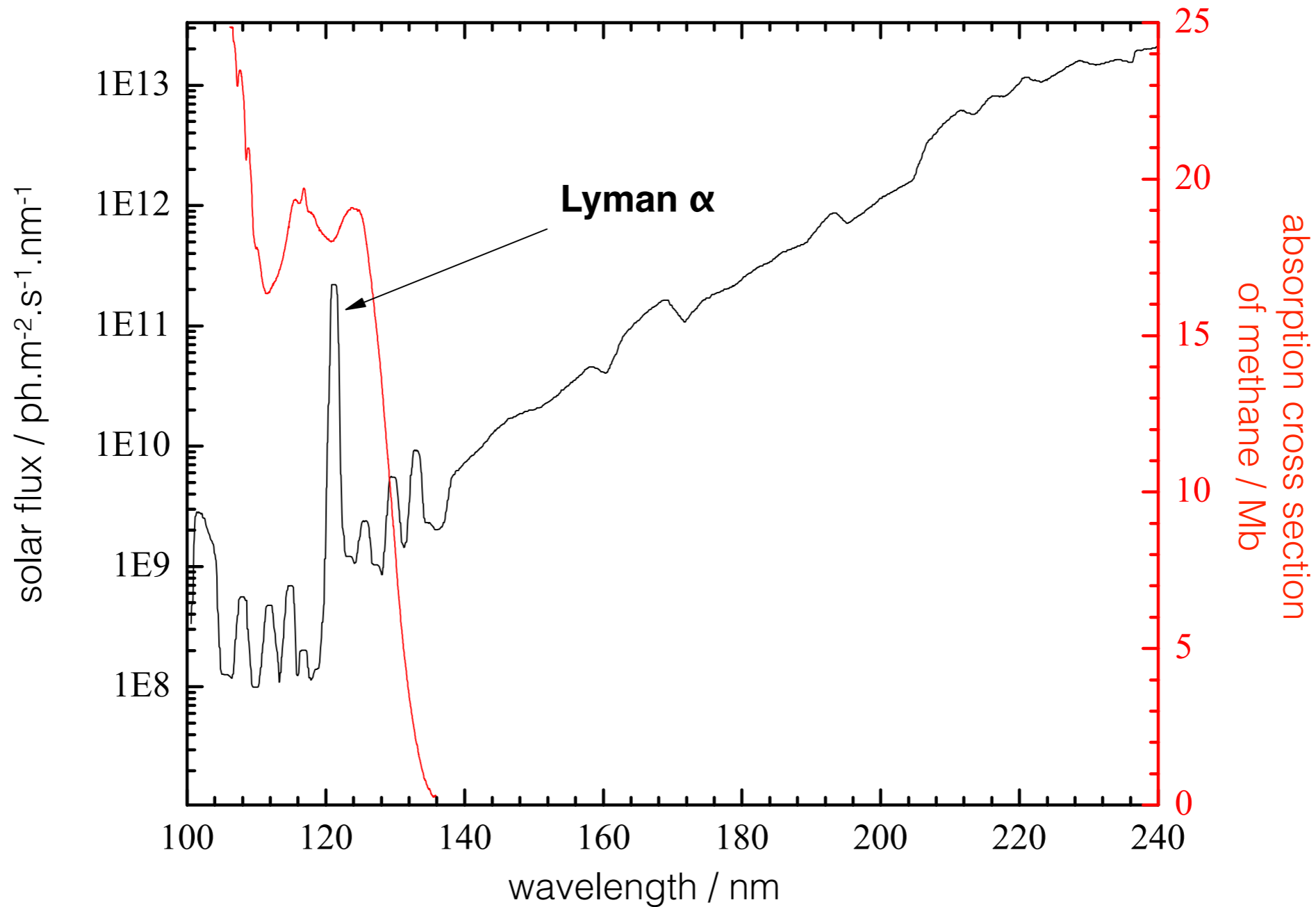
## Titan:

- Biggest moon of Saturn
- Dense atmosphere ( $P_s \approx 1.5$  bar)
- $T_s \approx 94$  K
- Composition:
  - $N_2 \approx 98$  %
  - $CH_4 \approx 1.6$  %
  - organic molecules
- Active atmosphere

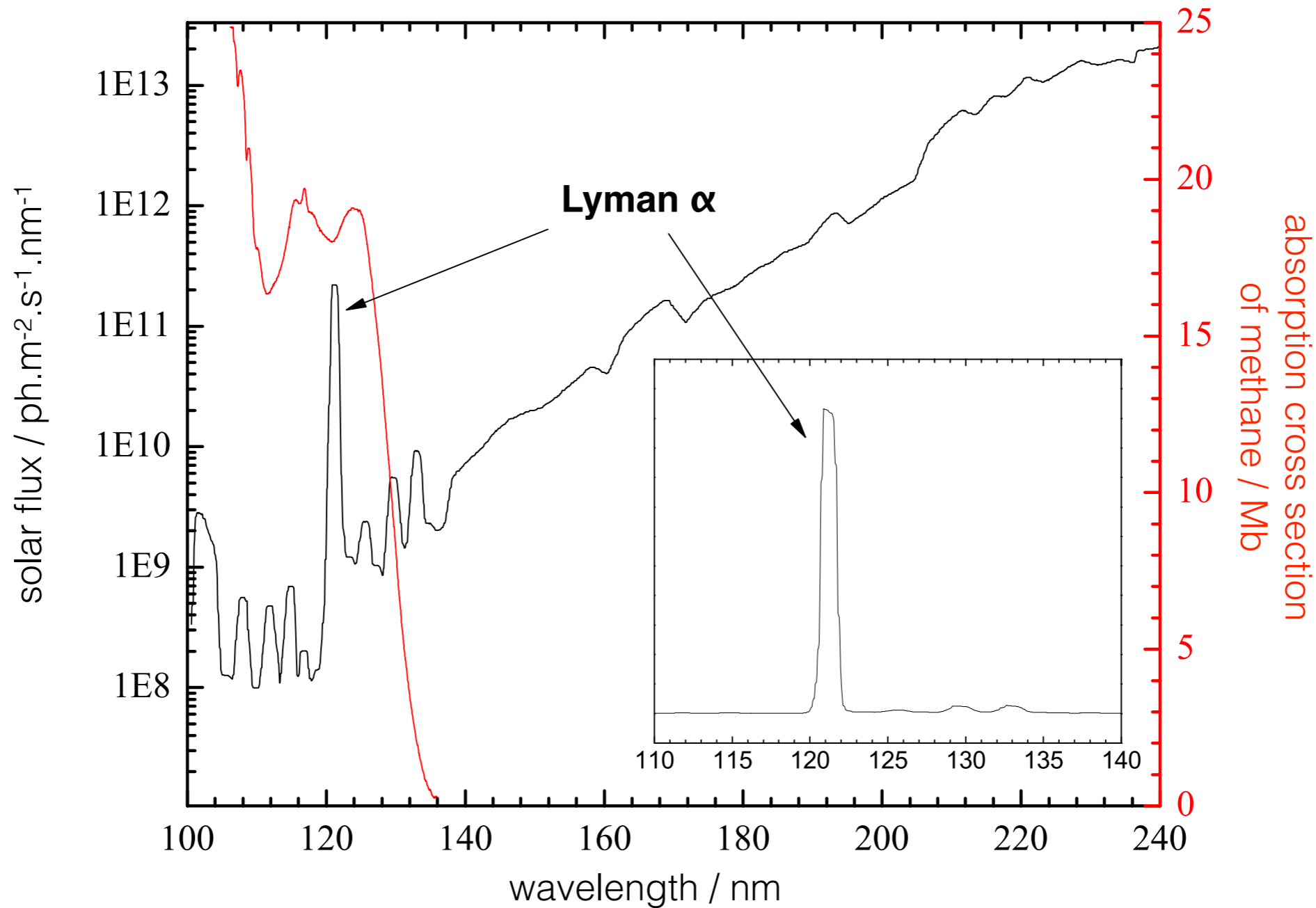
# Photochemistry on Titan



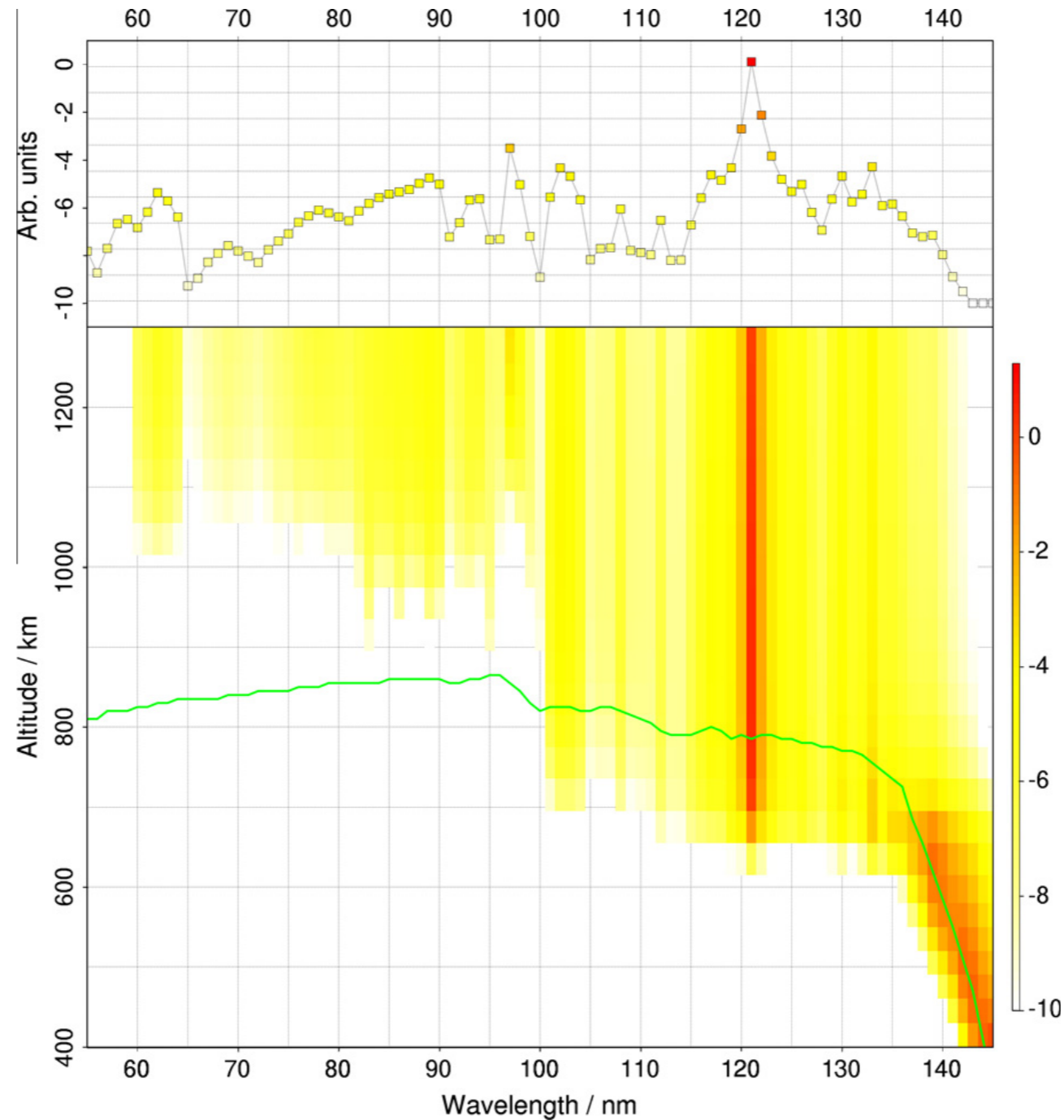
# Methane photolysis: a key process for Titan photochemistry



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# Methane photolysis: a key process for Titan photochemistry



B. Gans, Z. Peng, N. Carrasco, D. Gauyacq, S. Lebonnois, P. Pernot, *Icarus* **223** (2013) 330–343

# Previous measurements at Lyman $\alpha$ (121.6 nm)

Dissociative channel	Mordaunt et al. S1 & S2 (1993)		Heck et al. (1995)	Brownsword et al. (1997)	Wang et al. (2000)	Park et al. (2008)
CH <sub>3</sub> (X)+H	0.51	0.49	0.66	0.38	0.291	0.31
CH <sub>2</sub> (a)+H <sub>2</sub>	0.24	0	0.22	0.52	0.584	0.63
CH <sub>2</sub> (b)+H <sub>2</sub>	/	/	/	/	/	/
CH <sub>2</sub> (X)+2H	0.05	$\approx 0$	$\approx 0$	0.01	/	/
CH <sub>2</sub> (a)+2H	0.2	$\approx 0$	$\approx 0$		0.055	/
CH(X)+H+H <sub>2</sub>	0	0.51	0.11	0.08	0.07	0.059
C( <sup>1</sup> D)+2H <sub>2</sub>	0	0	0	0	0	0.0004
$\Phi$ (H)	1	1	0.77	0.47	0.47	0.31
$\Phi$ (H <sub>2</sub> )	0.24	0.51	0.33	0.6	0.654	0.69

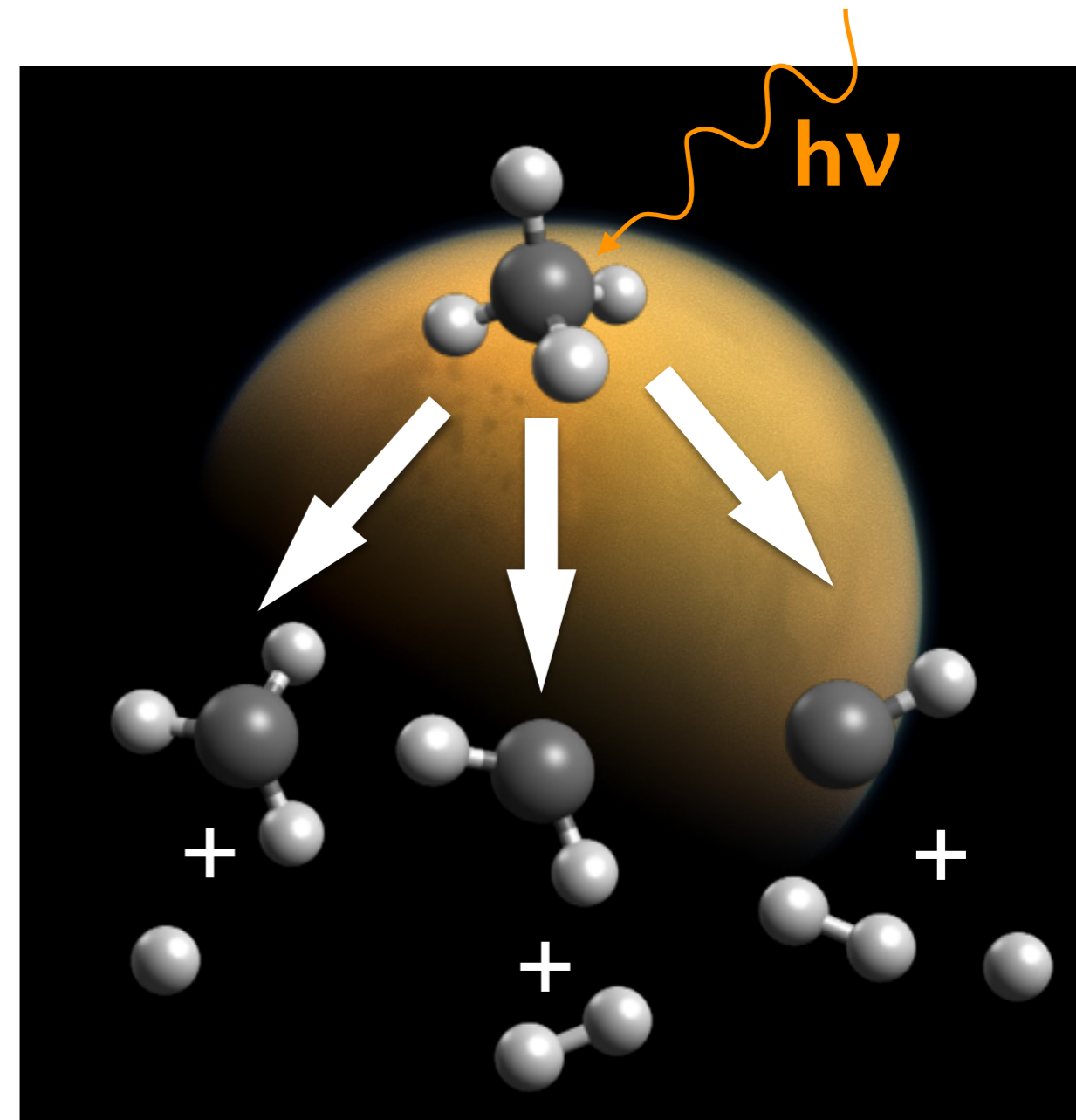


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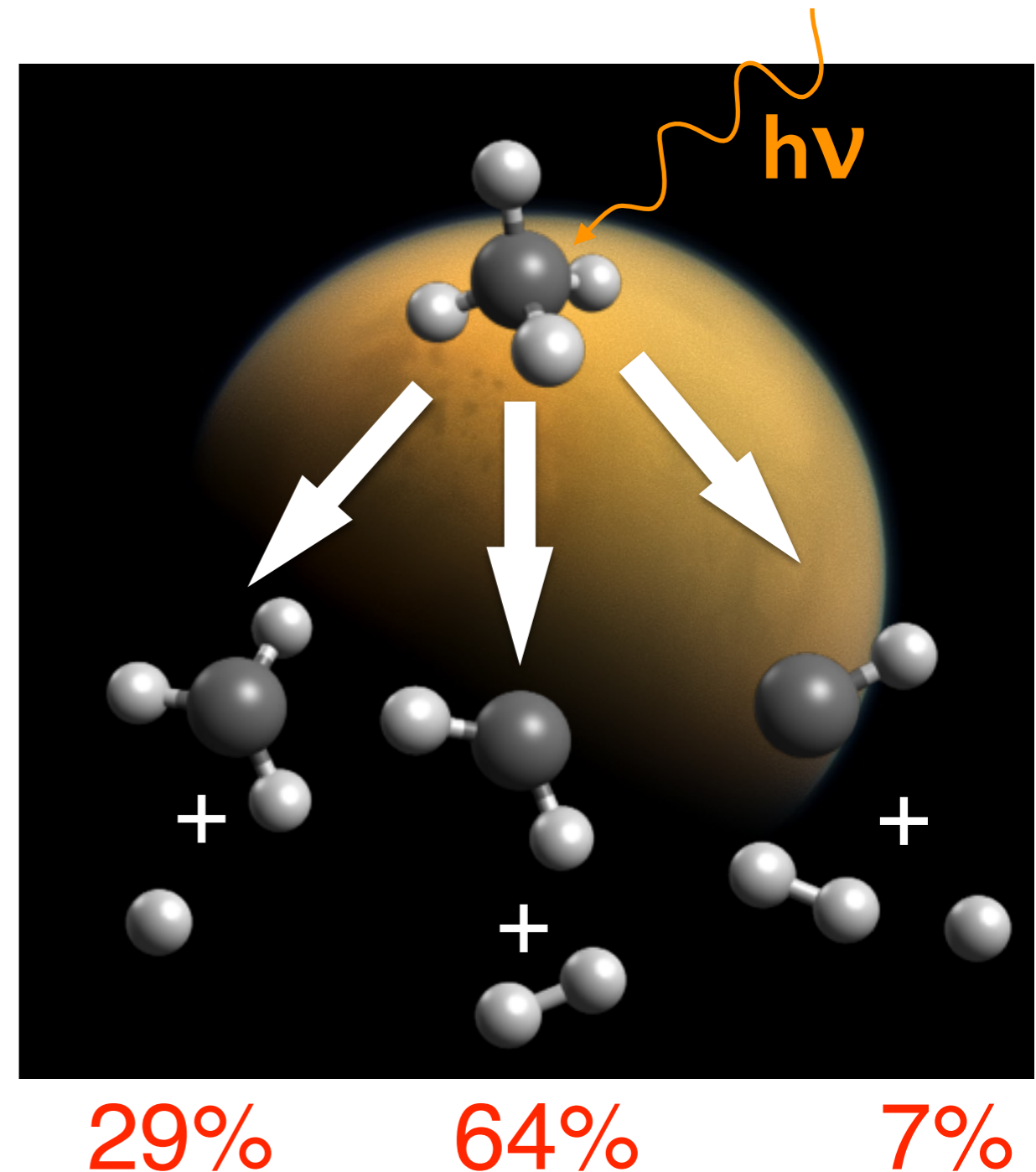
# How was methane photolysis implemented in the Titan atmosphere models until recently?

- 3 dissociation pathways:
  - $\text{CH}_3 + \text{H}$
  - $\text{CH}_2 + \text{H}_2$
  - $\text{CH} + \text{H} + \text{H}_2$
- Two different energy ranges:
  - at Lyman- $\alpha$  (121.6 nm)
  - Out Lyman- $\alpha$



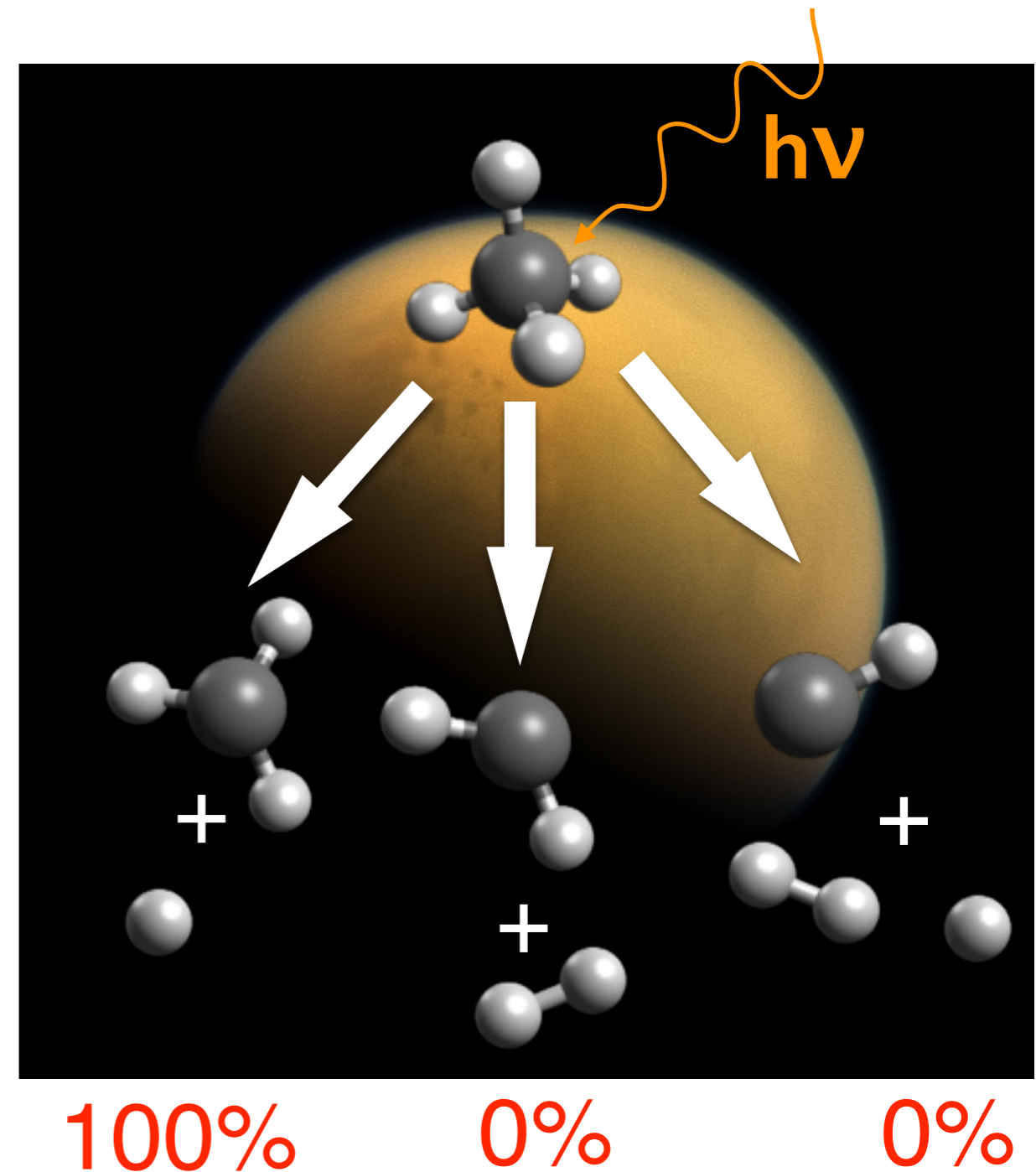
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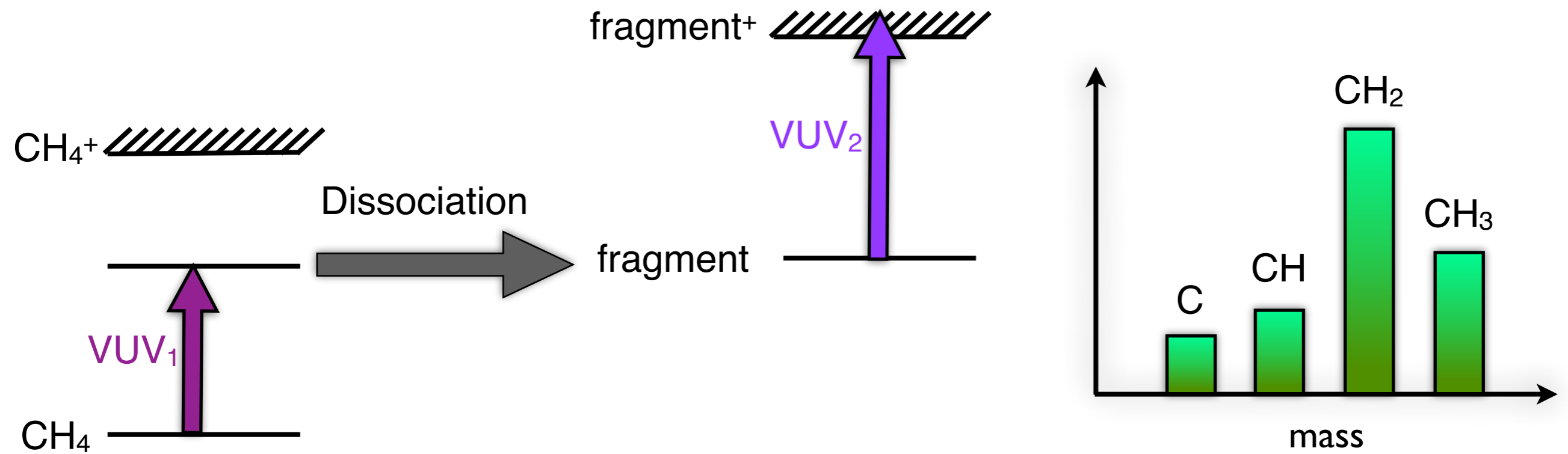


# What is important?

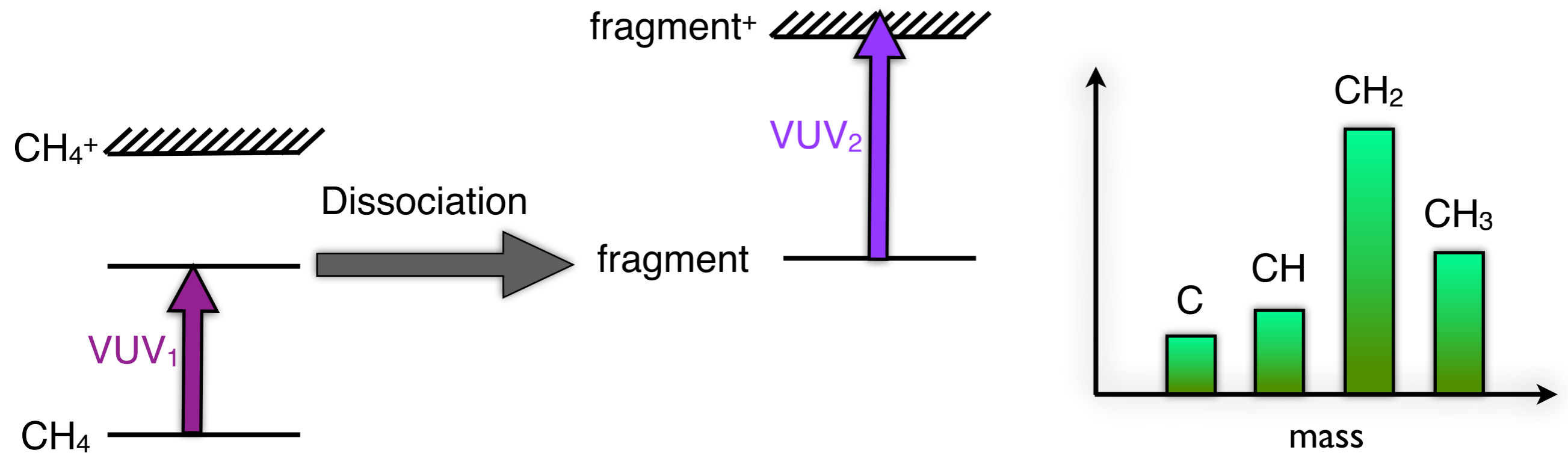
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- from the simulations:
  - **CH<sub>2</sub>** and **CH<sub>3</sub>** are the most important products,
  - **Lyman  $\alpha$**  is the most influent wavelength, but «out-Lyman  $\alpha$ » wavelengths have an effect,
- from the literature:
  - **direct probing of CH<sub>2</sub> and CH<sub>3</sub> fragments is mandatory**

# Strategy

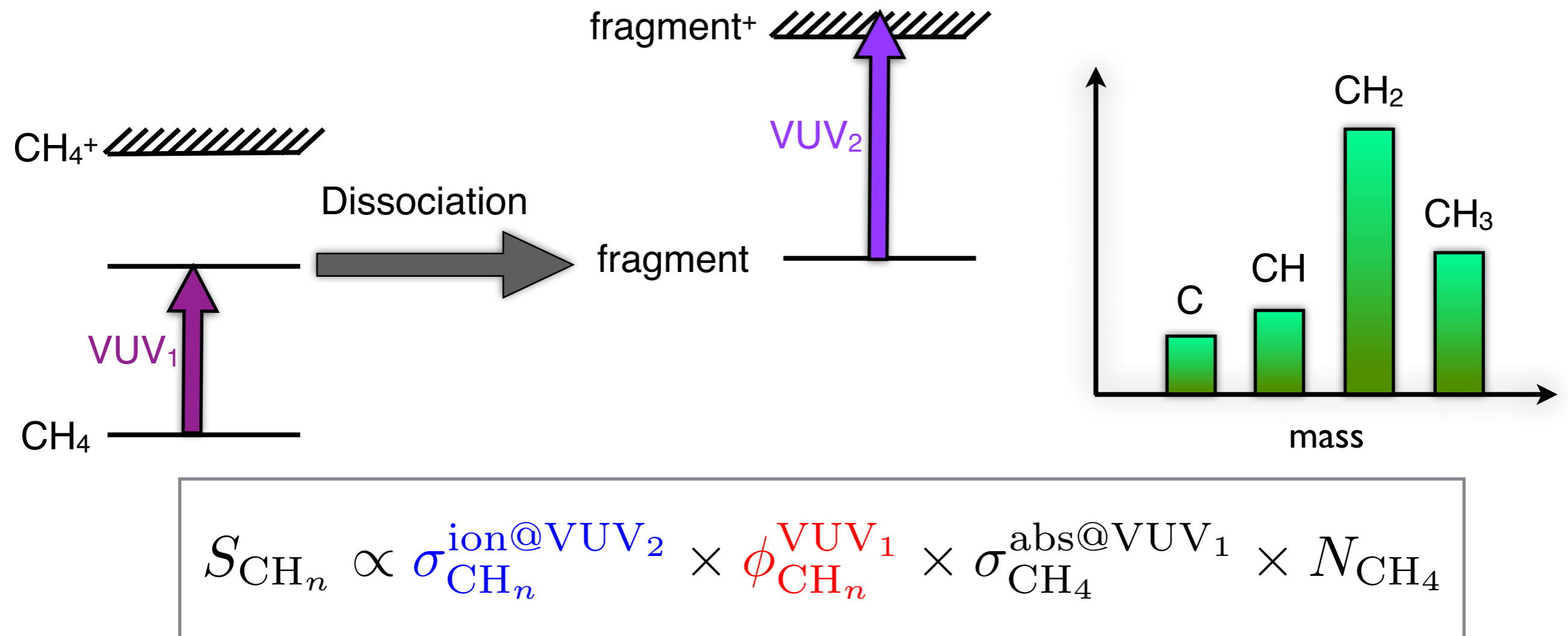


# Strategy



$$S_{\text{CH}_n} \propto \sigma_{\text{CH}_n}^{\text{ion@VUV}_2} \times \phi_{\text{CH}_n}^{\text{VUV}_1} \times \sigma_{\text{CH}_4}^{\text{abs@VUV}_1} \times N_{\text{CH}_4}$$

# Strategy



➔ **Branching ratios** can be extracted from the mass spectrum if the **photoionization cross sections of the radical products** are known

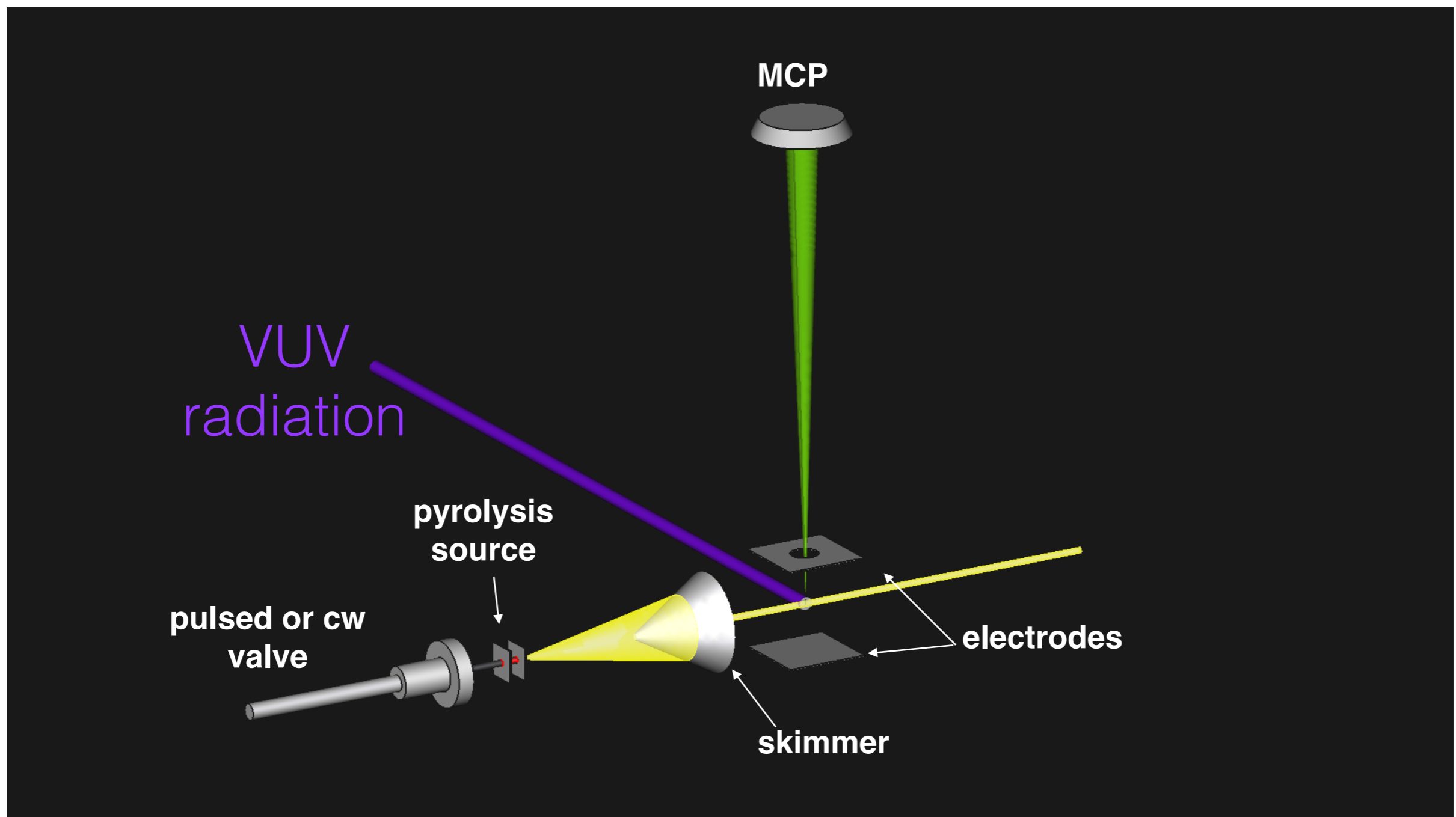


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First step: absolute photoionization  
cross section measurements of radicals

---

# Experimental setup for radical cross section measurements

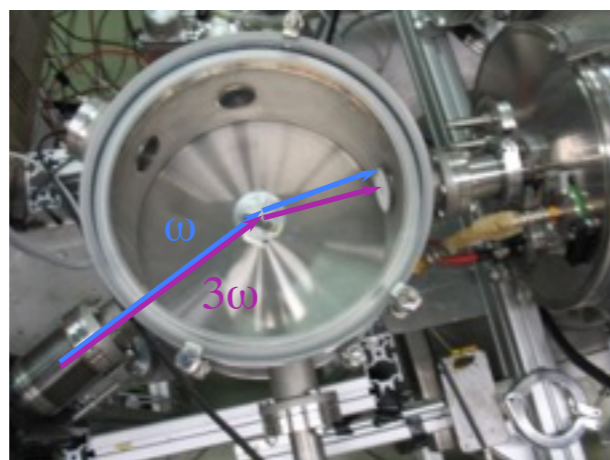
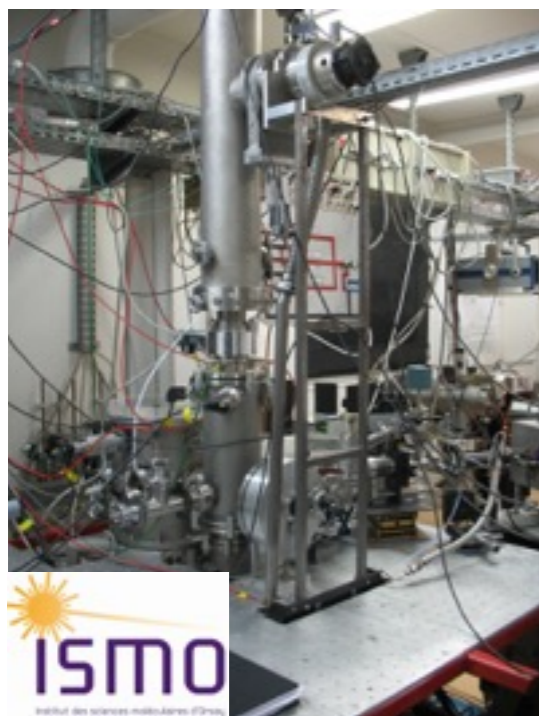


# VUV sources

In the lab (ISMO)

VUV laser at 118.2 nm ( $3\omega$ )

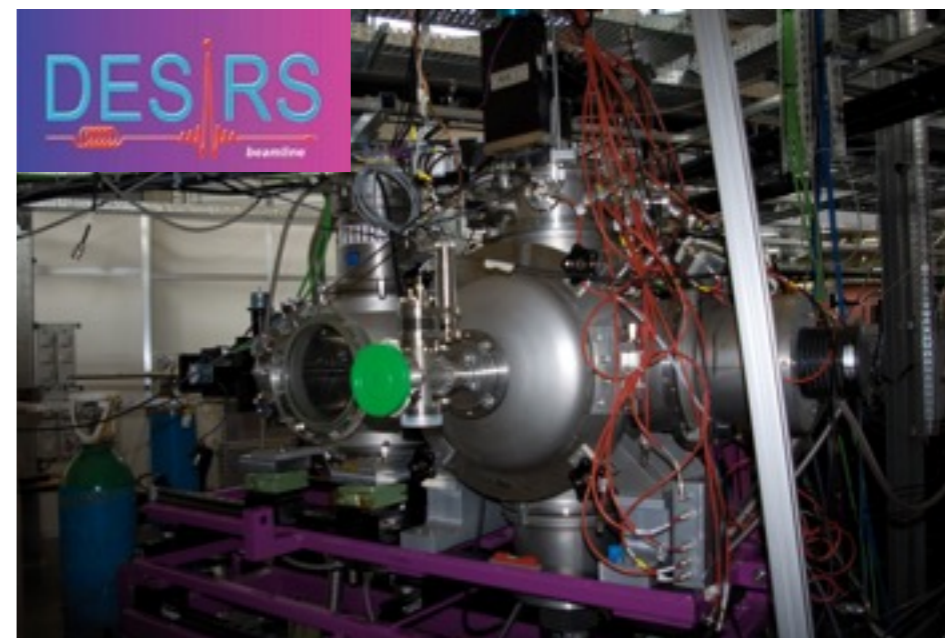
(84602.4  $\text{cm}^{-1}$ )



SOLEIL synchrotron

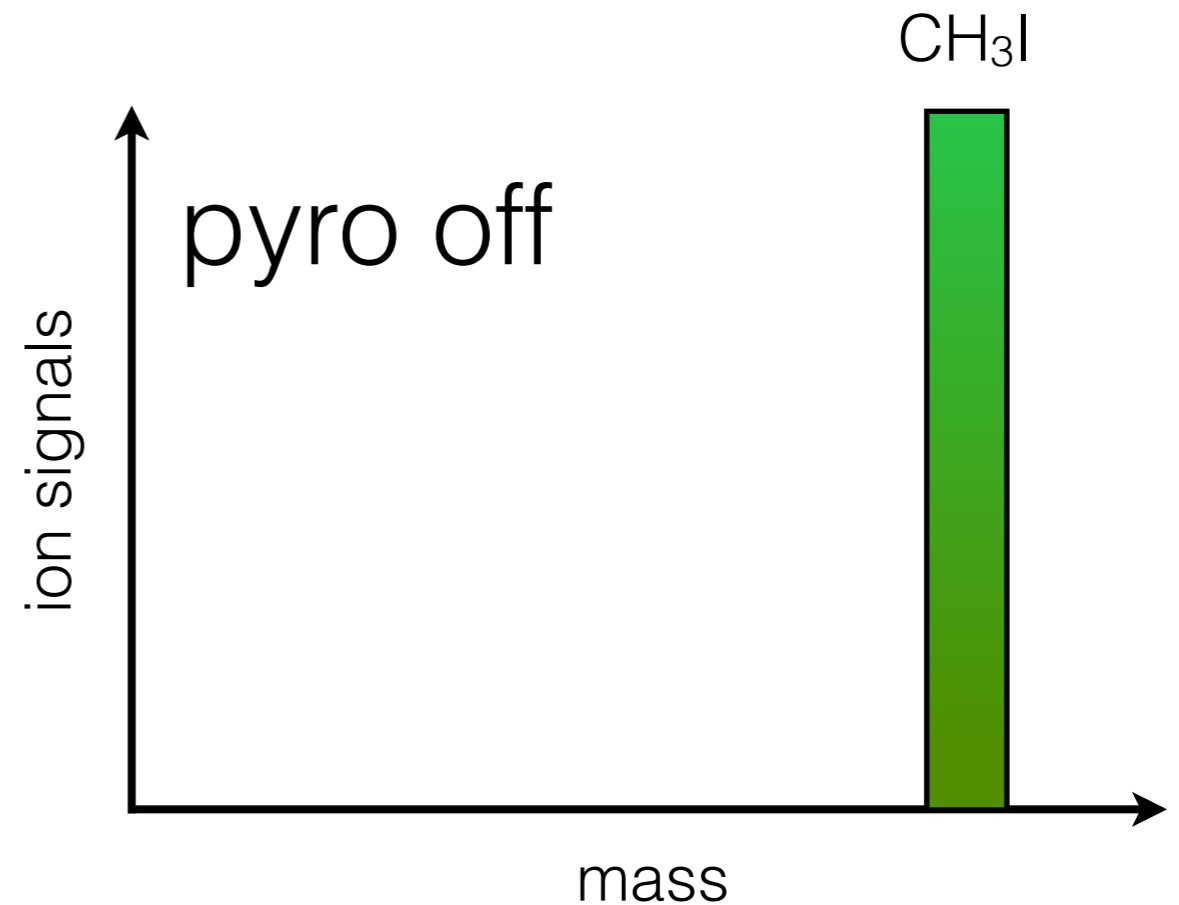
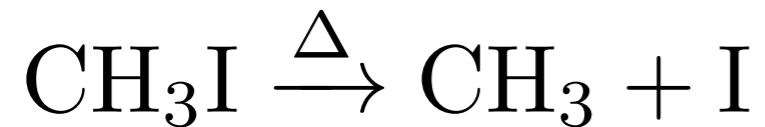
8  $\rightarrow$  12 eV

(64524.4  $\rightarrow$  96786.5  $\text{cm}^{-1}$ )



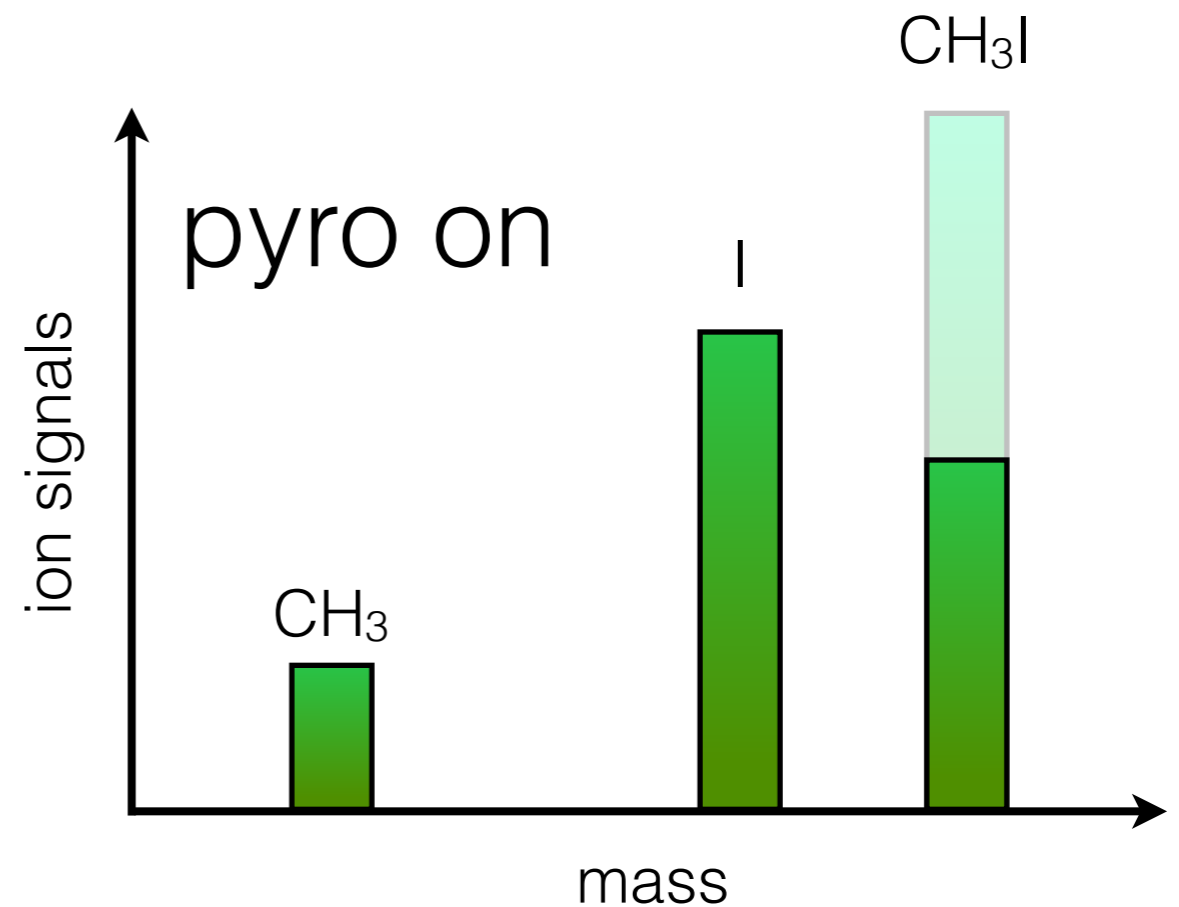
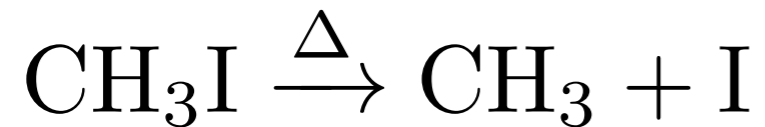
# Method for measuring absolute photoionization cross sections

Pyrolysis of CH<sub>3</sub>I :



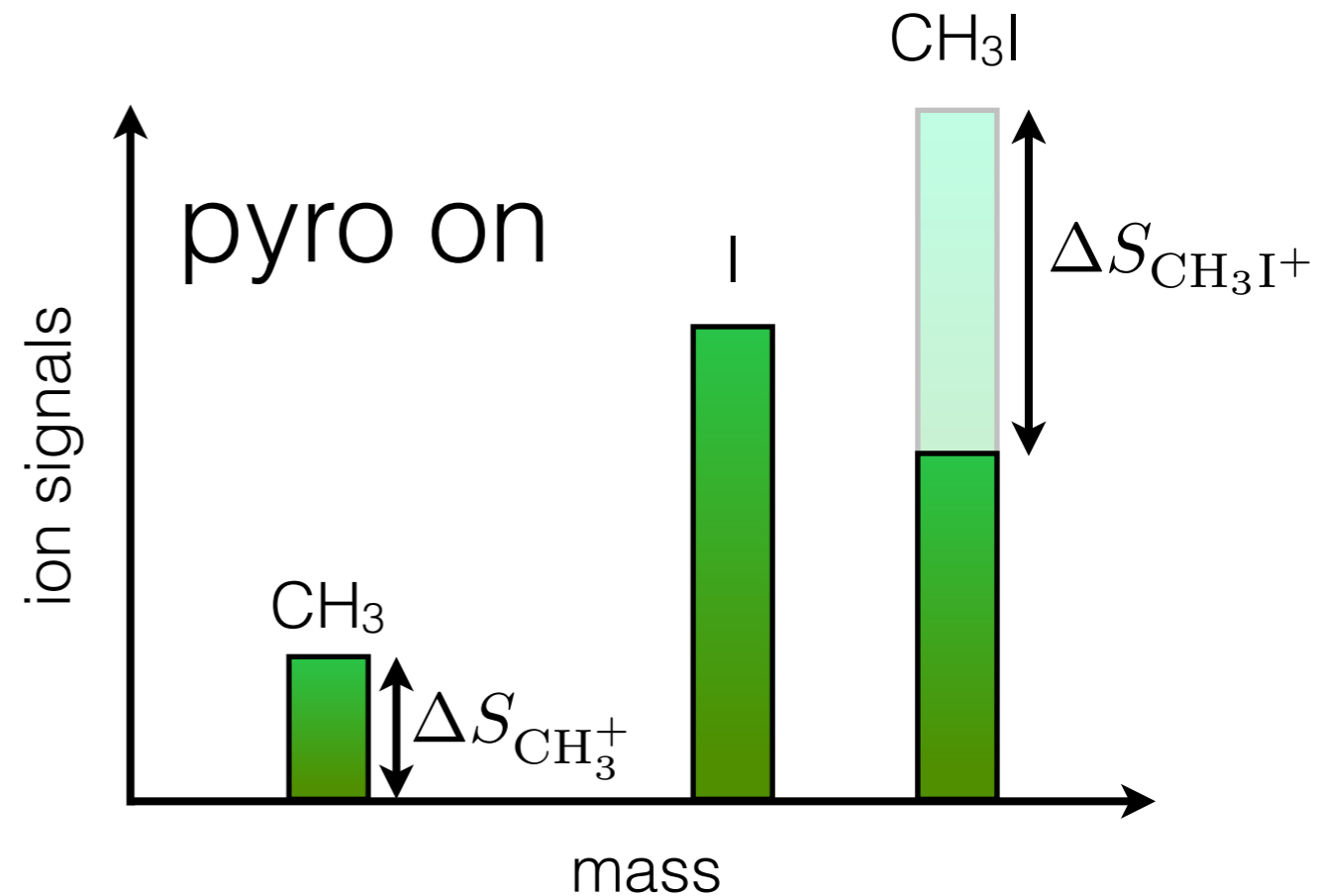
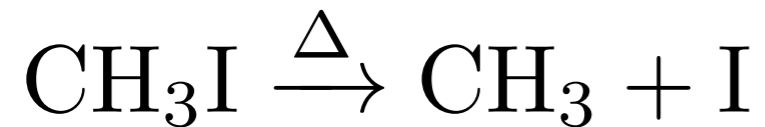
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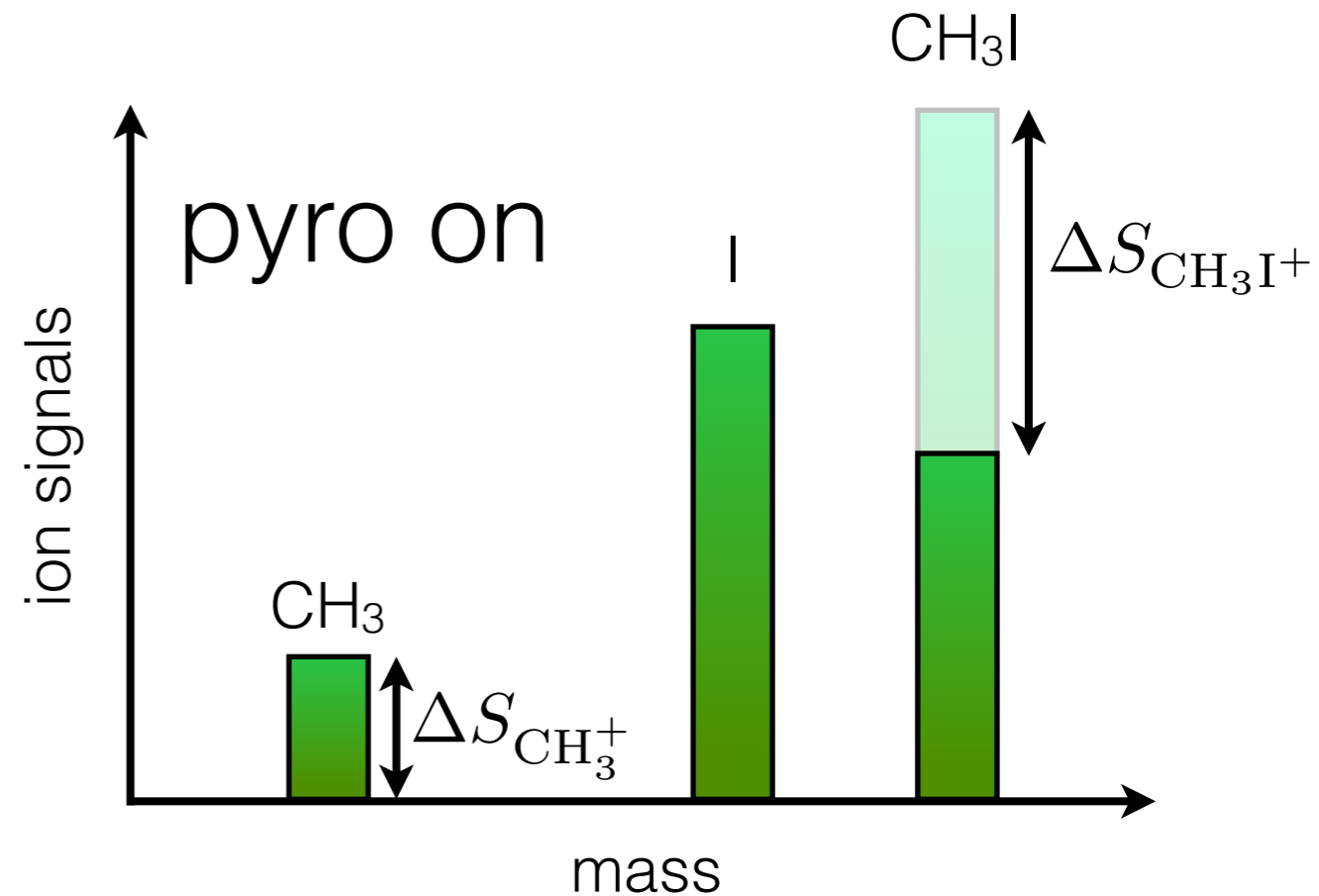
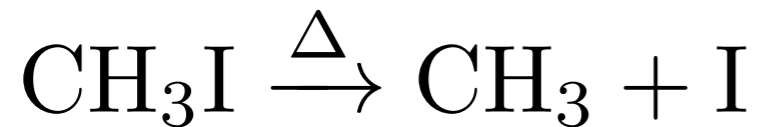
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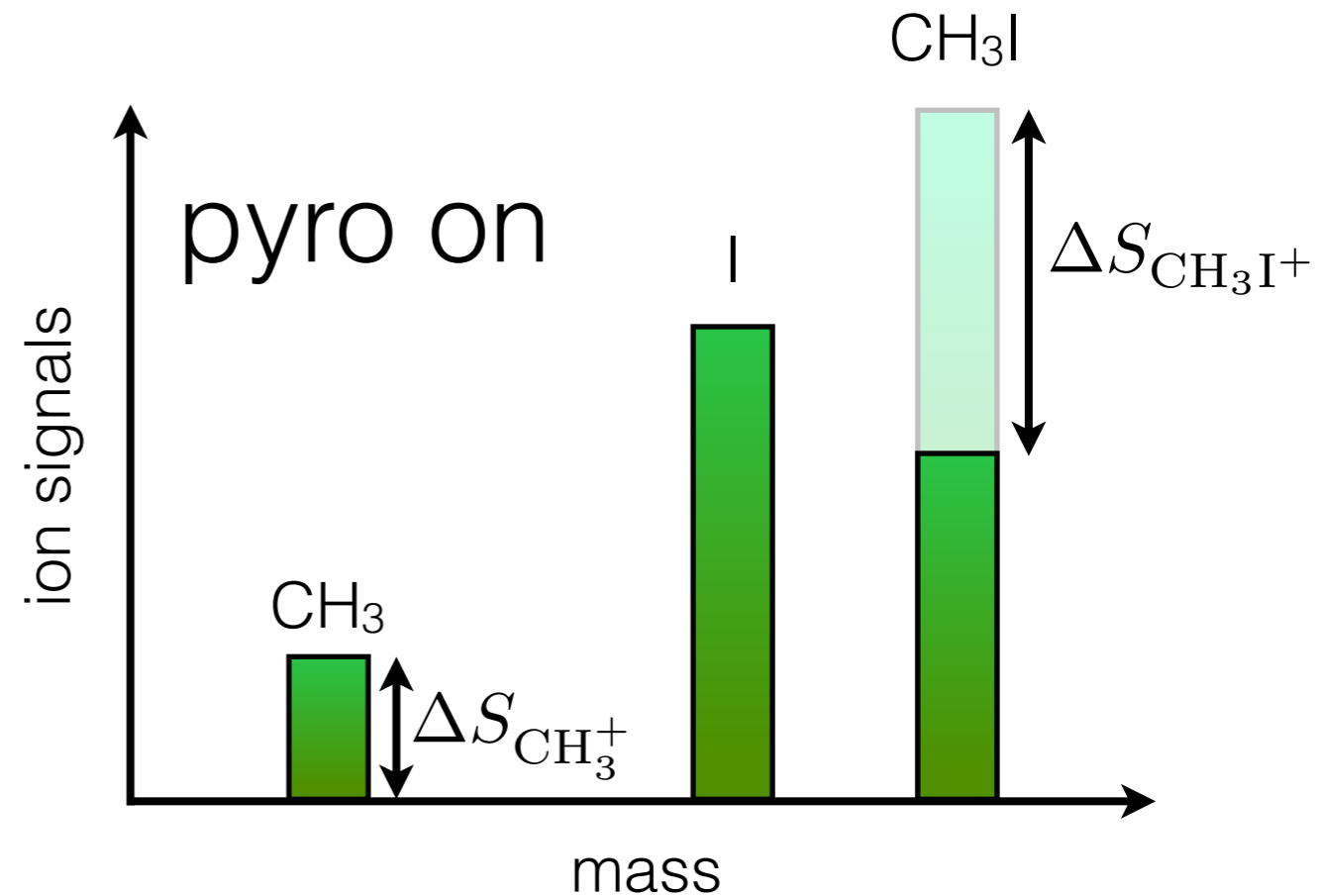
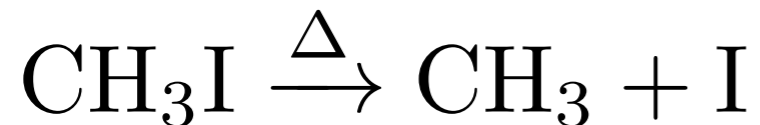
Pyrolysis of CH<sub>3</sub>I :



$$\frac{\Delta S_{\text{CH}_3^+}}{\Delta S_{\text{CH}_3\text{I}^+}} \times \frac{F_{\text{CH}_3\text{I}}}{F_{\text{CH}_3}} = \frac{\sigma_{\text{CH}_3}^{i@118,2}}{\sigma_{\text{CH}_3\text{I}}^{i@118,2}}$$

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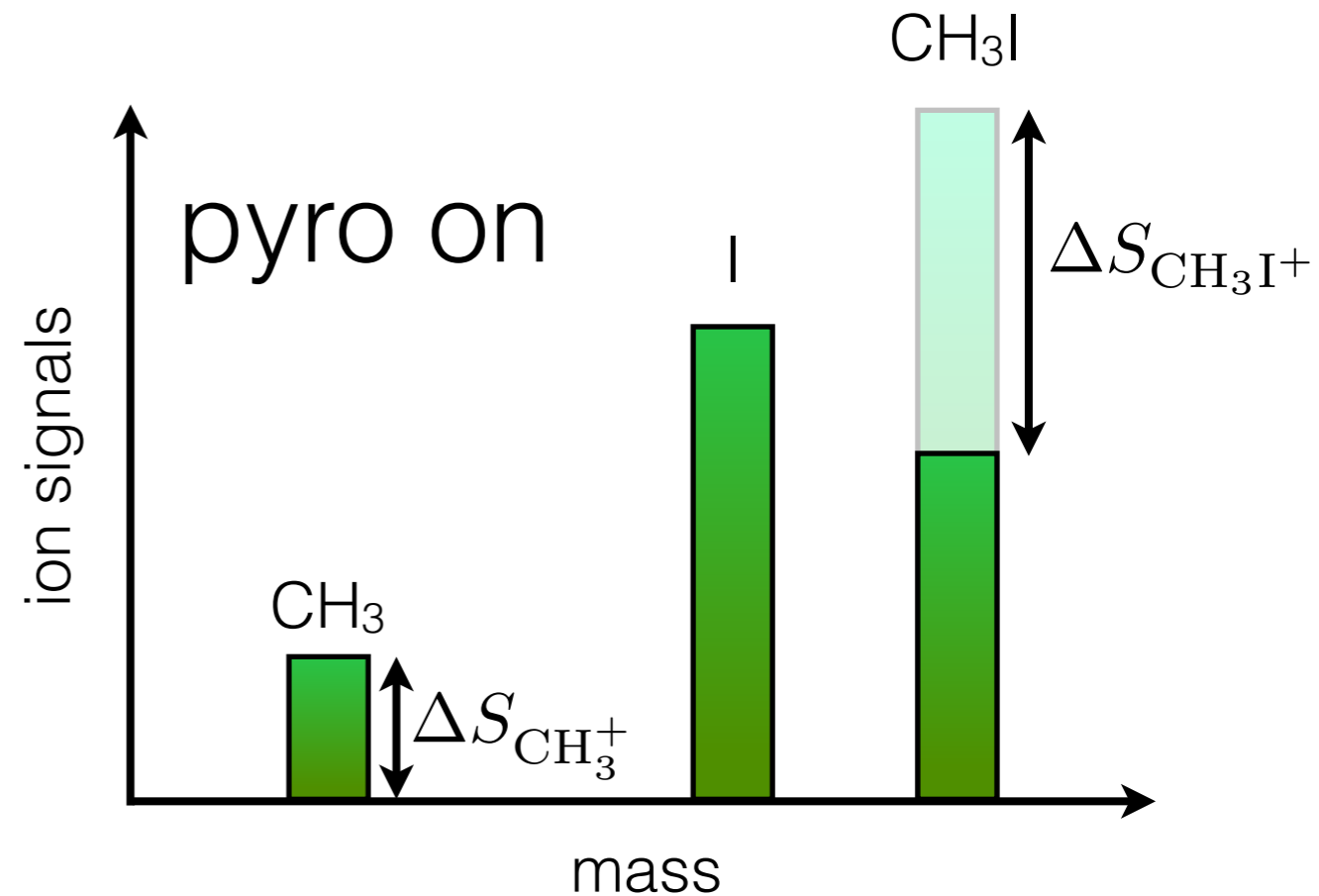
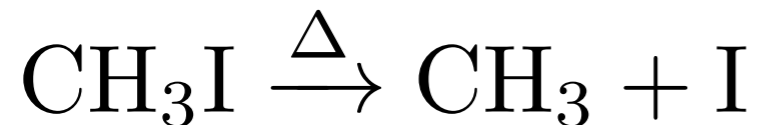


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# How to measure F mass discrimination factor?

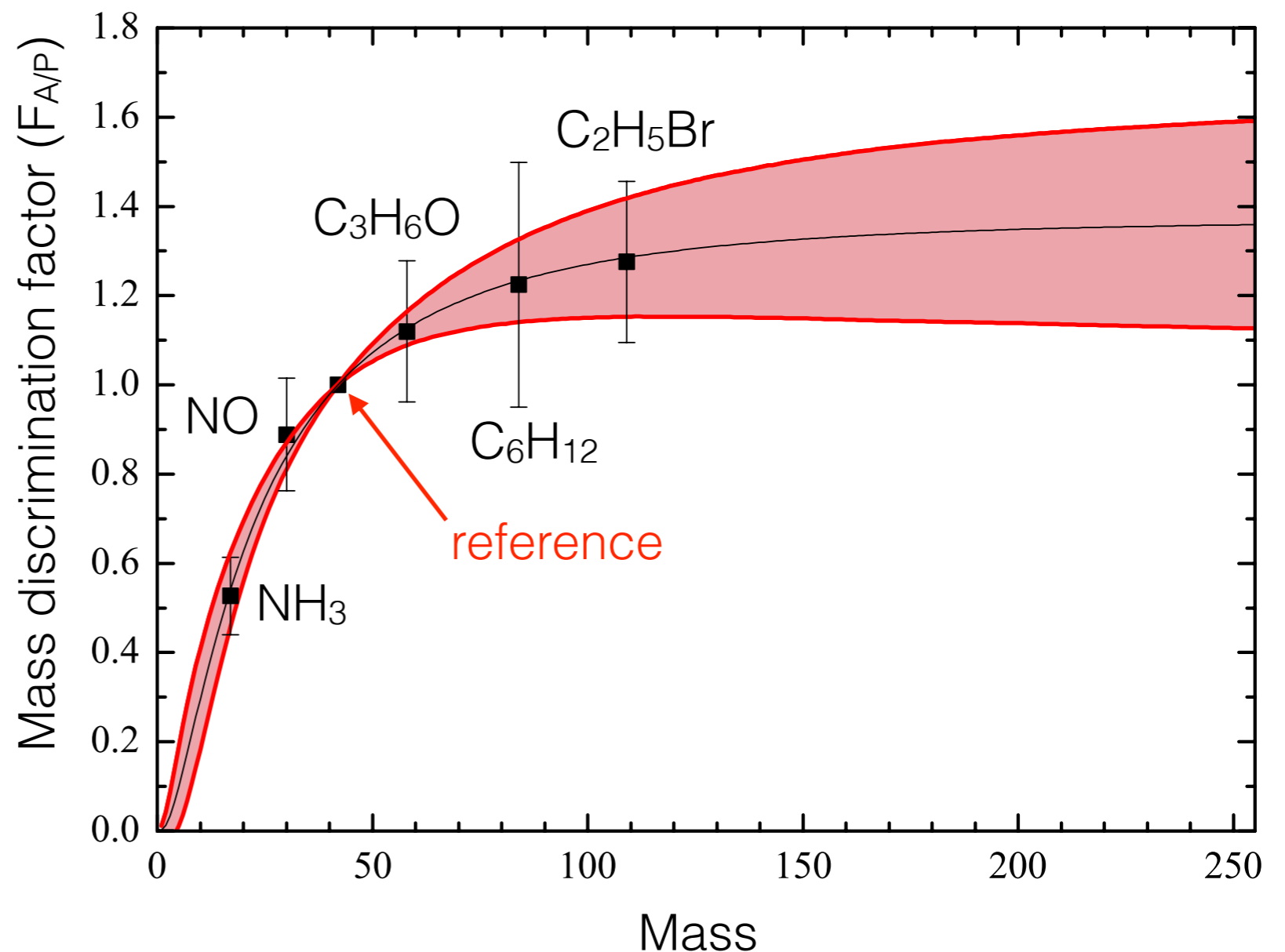
## - 6 molecules :

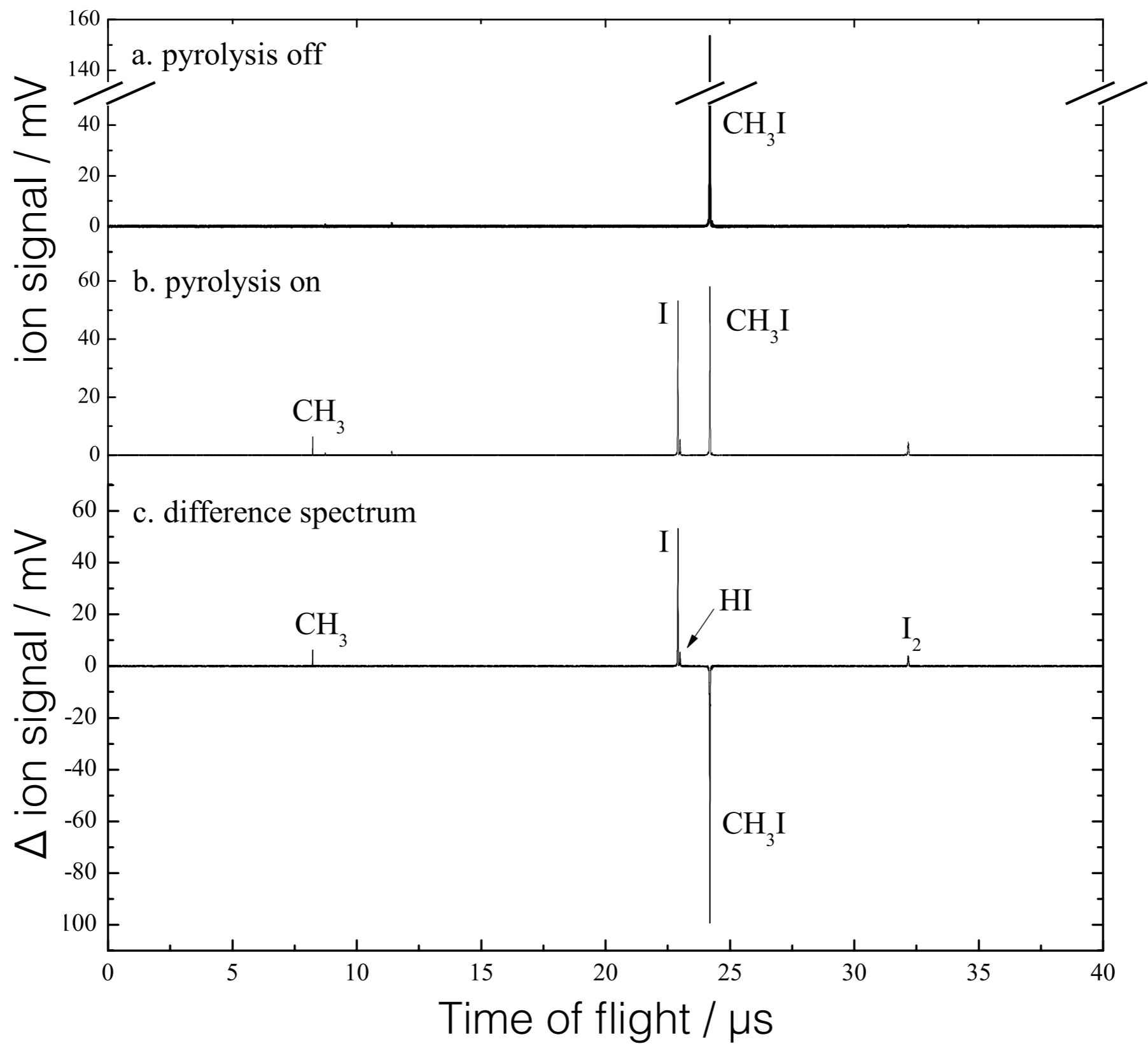
- $\text{NH}_3$  (Ammonia)
- $\text{NO}$  (Nitric oxide)
- $\text{C}_3\text{H}_6$  (propene, reference)
- $\text{C}_3\text{H}_6\text{O}$  (acetone)
- $\text{C}_6\text{H}_{12}$  (cyclohexane)
- $\text{C}_2\text{H}_5\text{Br}$  (bromoethane)

## - Extrapolation

(collaboration with P. Pernot, LCP)

- best fit
- uncertainty propagation





# Use of pyrolysis radical source ...under control!

---

- First requirement: Photoionization cross section independent with temperature (in the pyrolysis temperature range).
- Corrections are needed:
  - Pyrolysis temperature has physical effects on the beam (**expansion facteur**),
  - **Reactivity** has to be controlled.

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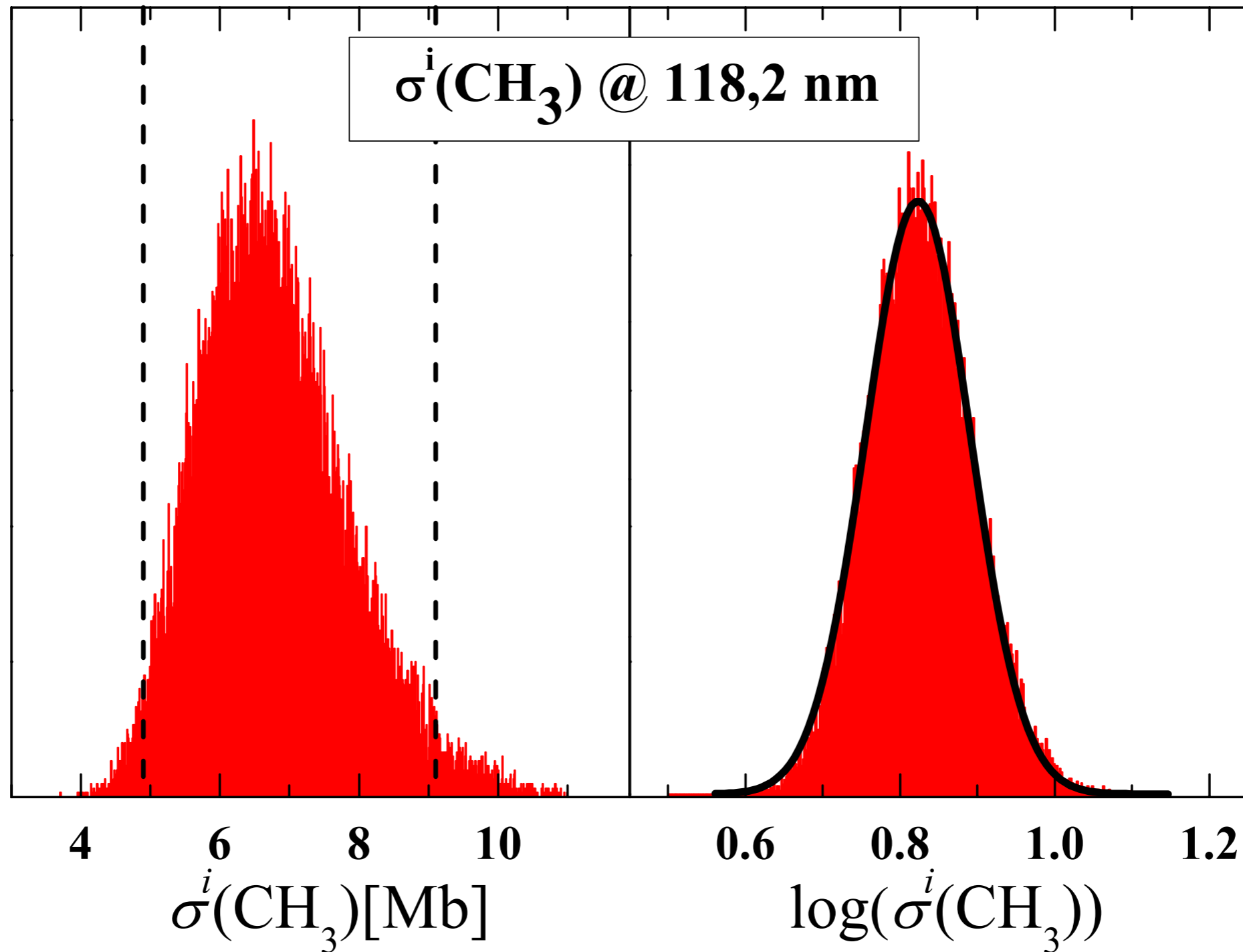
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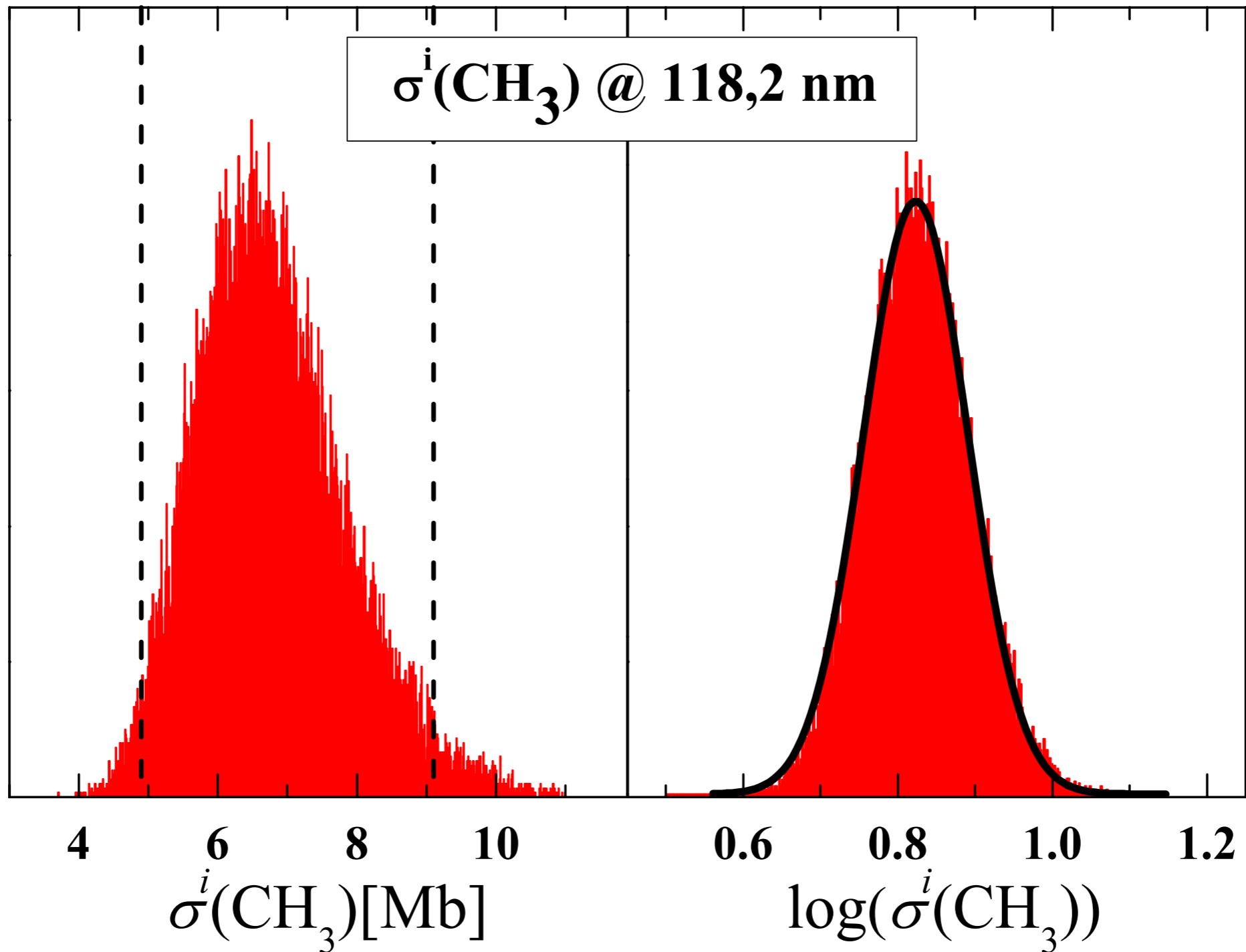
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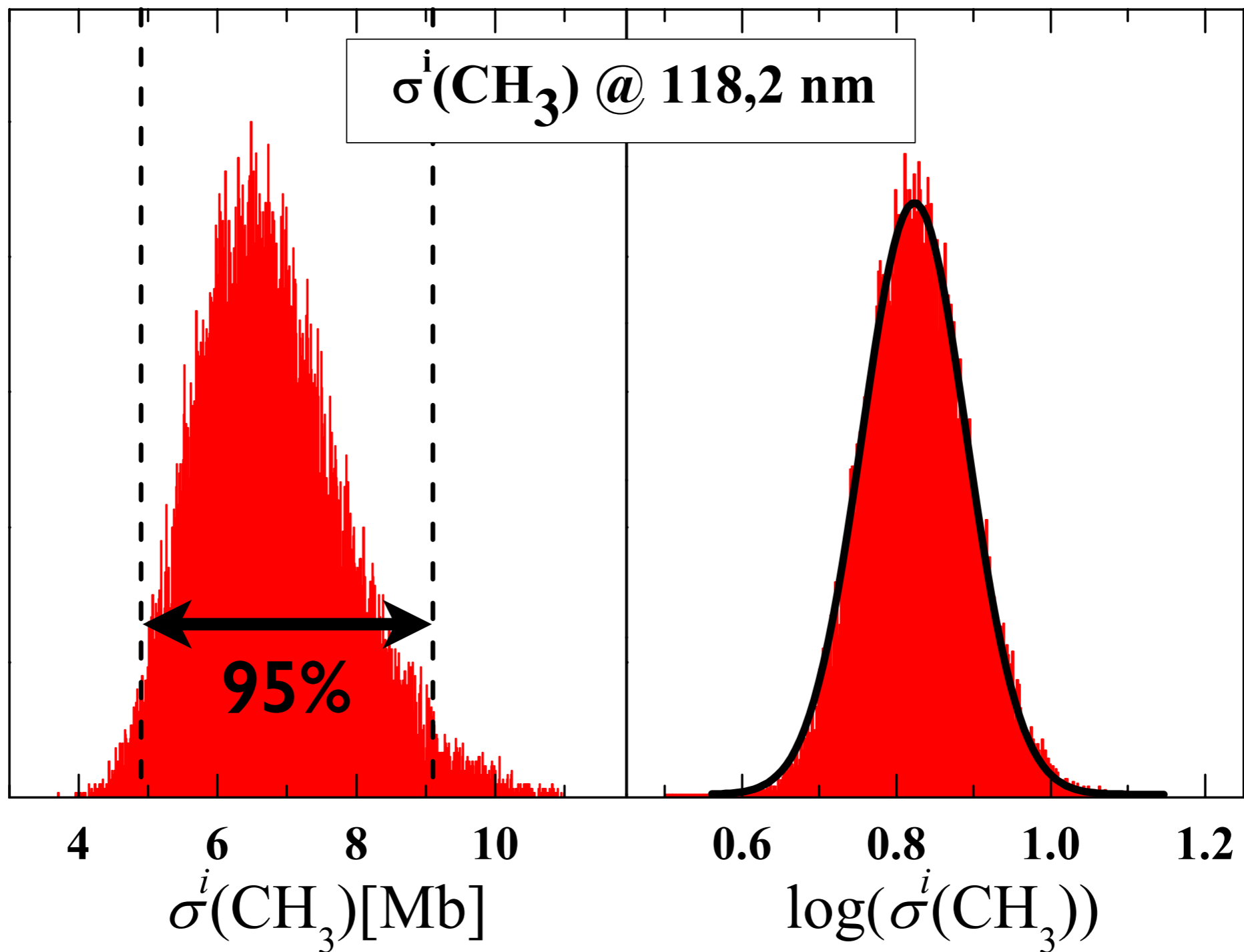
# Photoionization cross section of CH<sub>3</sub> at 118.2 nm



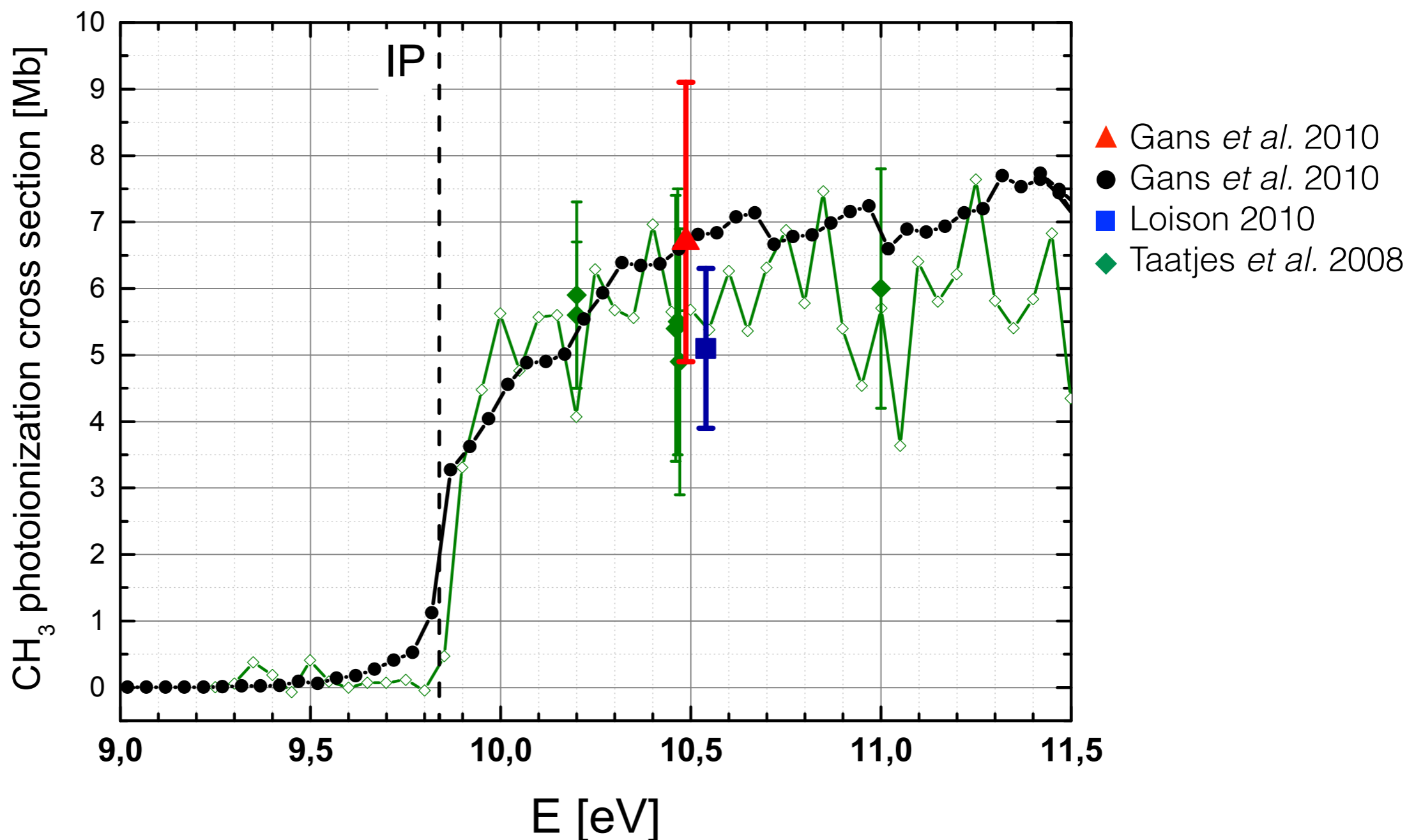
$$\sigma_{\text{CH}_3}^{i@118,2} = 6,7^{+2,4}_{-1,8} \text{ Mb}$$



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# Energy-dependent photoionization cross section: Combination Laser / synchrotron experiments



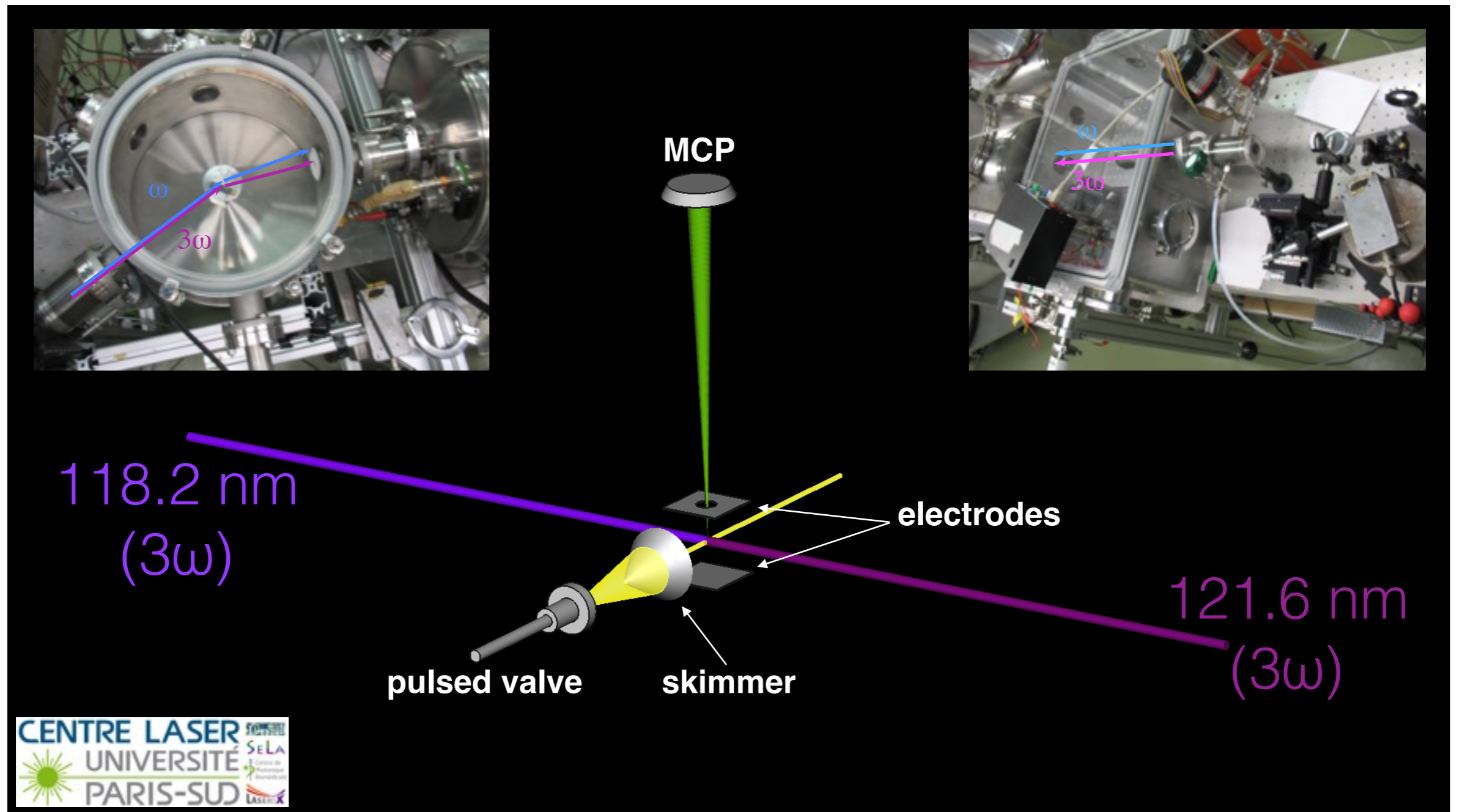
B. Gans, L. A. Vieira Mendes, S. Boyé-Péronne, S. Douin, G. Garcia, H. Soldi-Lose, B. K. Cunha de Miranda, C. Alcaraz, N. Carrasco, P. Pernot and D. Gauyacq, *J. Phys. Chem. A* **114**, 3237–3246 (2010)

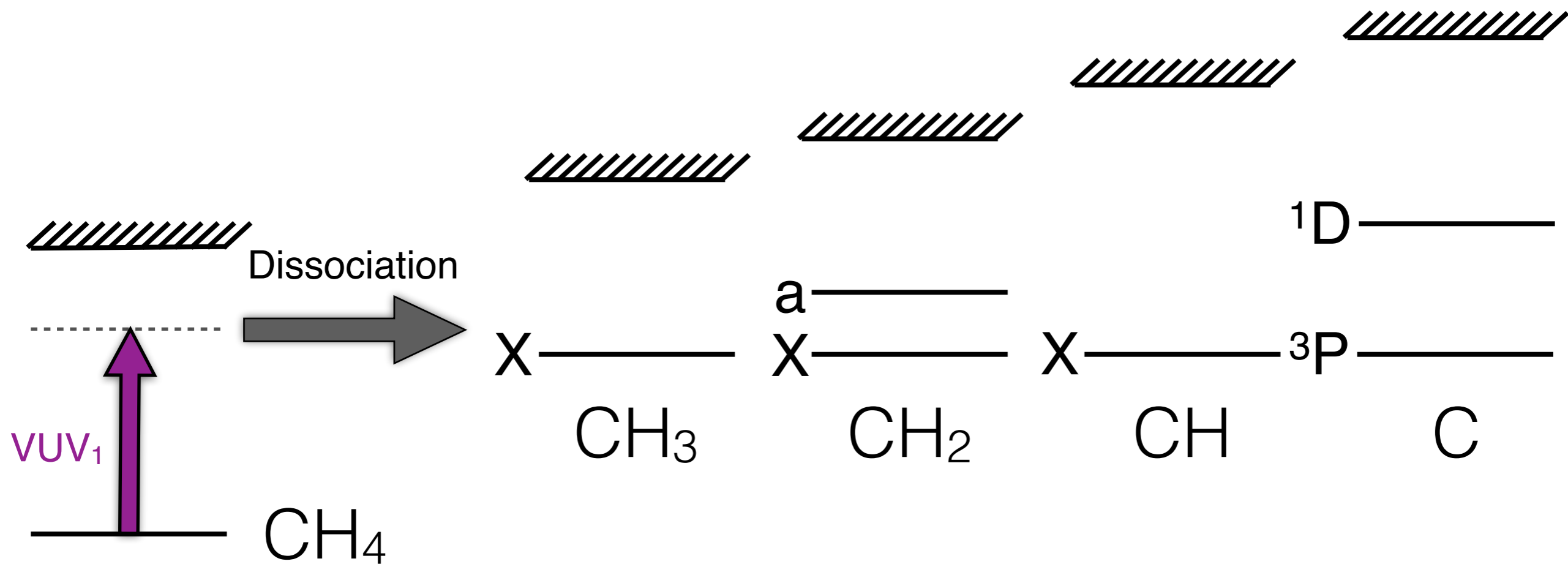
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# Second step: Methane photolysis experiment

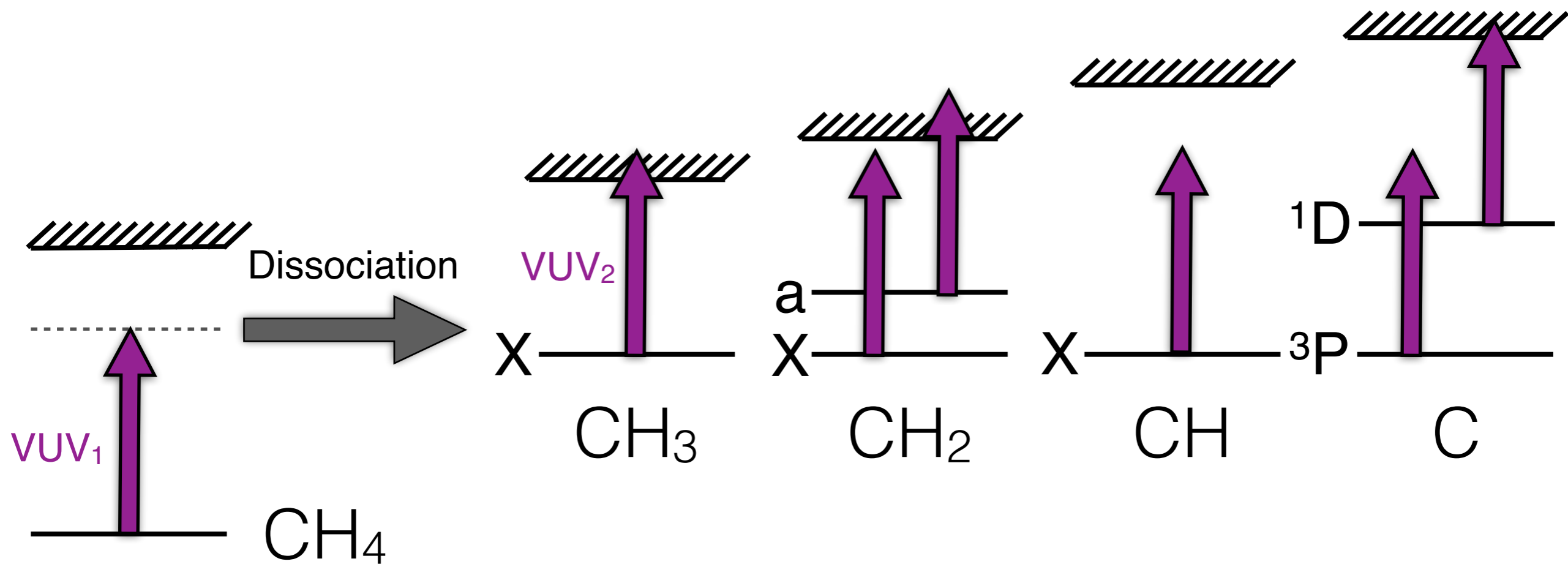
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# Experimental setup at CLUPS laser center: a VUV pump-VUV probe laser experiment



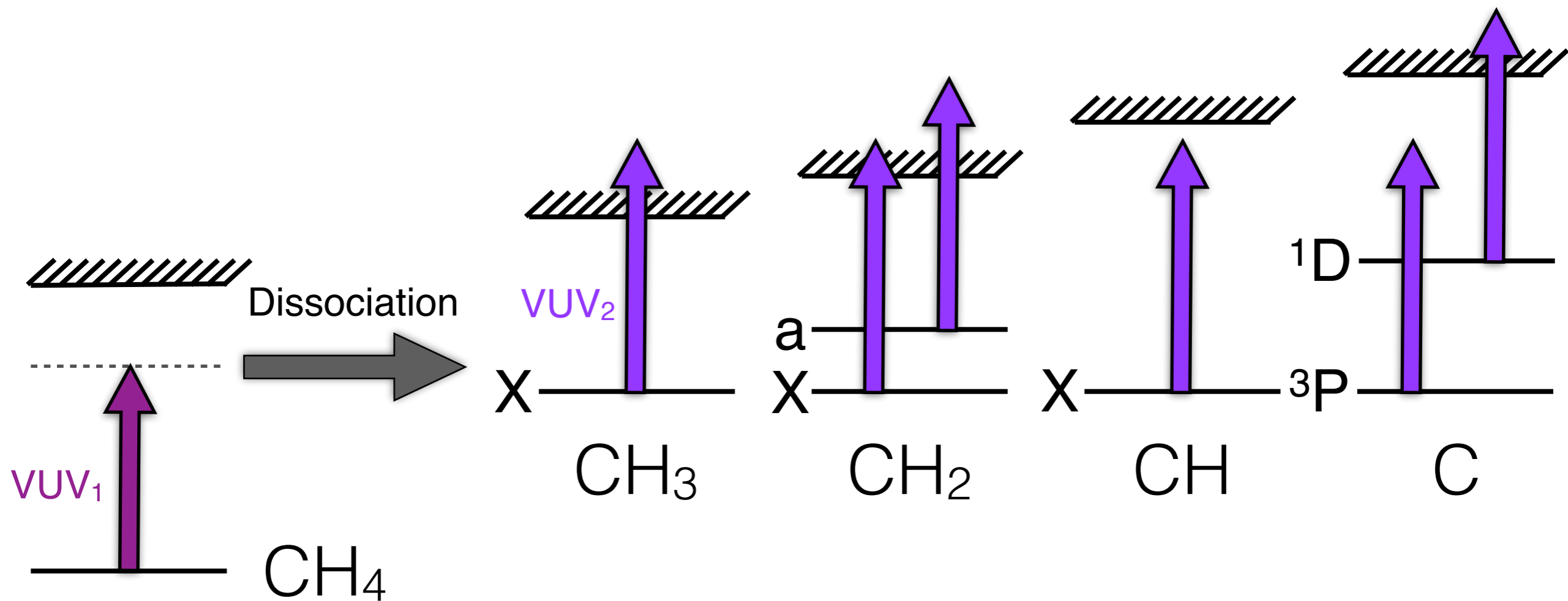


	$\text{CH}_3(\text{X})$	$\text{CH}_2(\text{X})$	$\text{CH}_2(\text{a})$	$\text{CH}(\text{X})$	$\text{C}(^1\text{D})$
IP [eV]	9.84	10.40	10.00	10.64	10.02
Ionization at 121.6 nm (10.2 eV)	✓	✗	✓	✗	✓
Ionization at 118.2 nm (10.49 eV)	✓	✓	✓	✗	✓



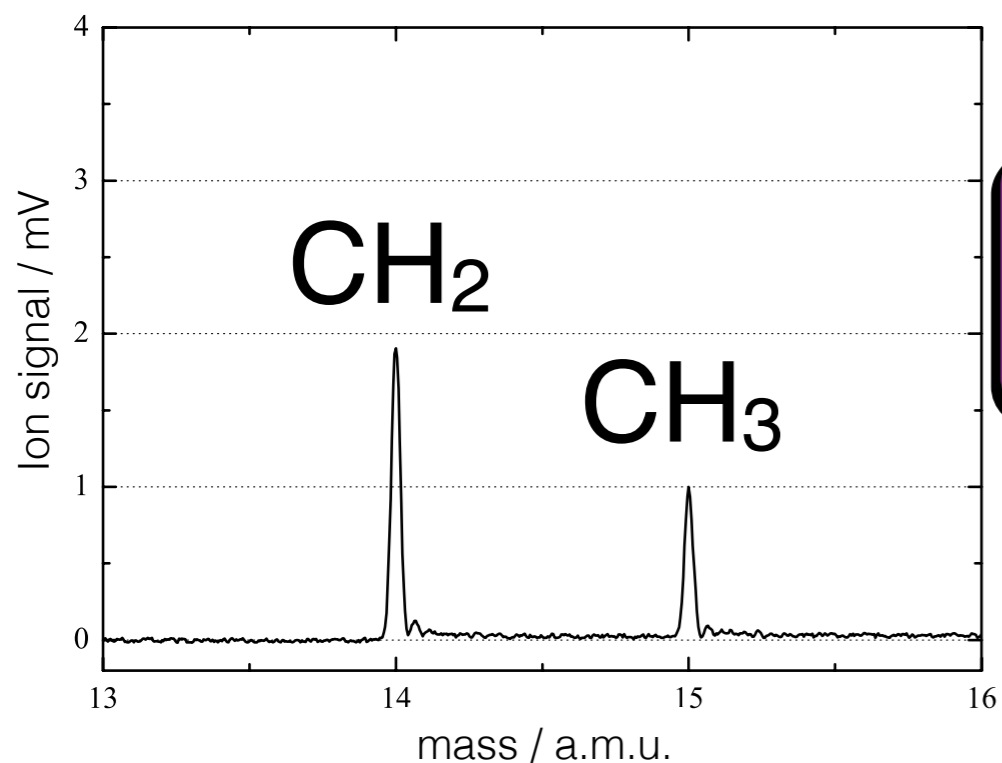
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Ionization at 118.2 nm (10.49 eV)	✓	✓	✓	✗	✓



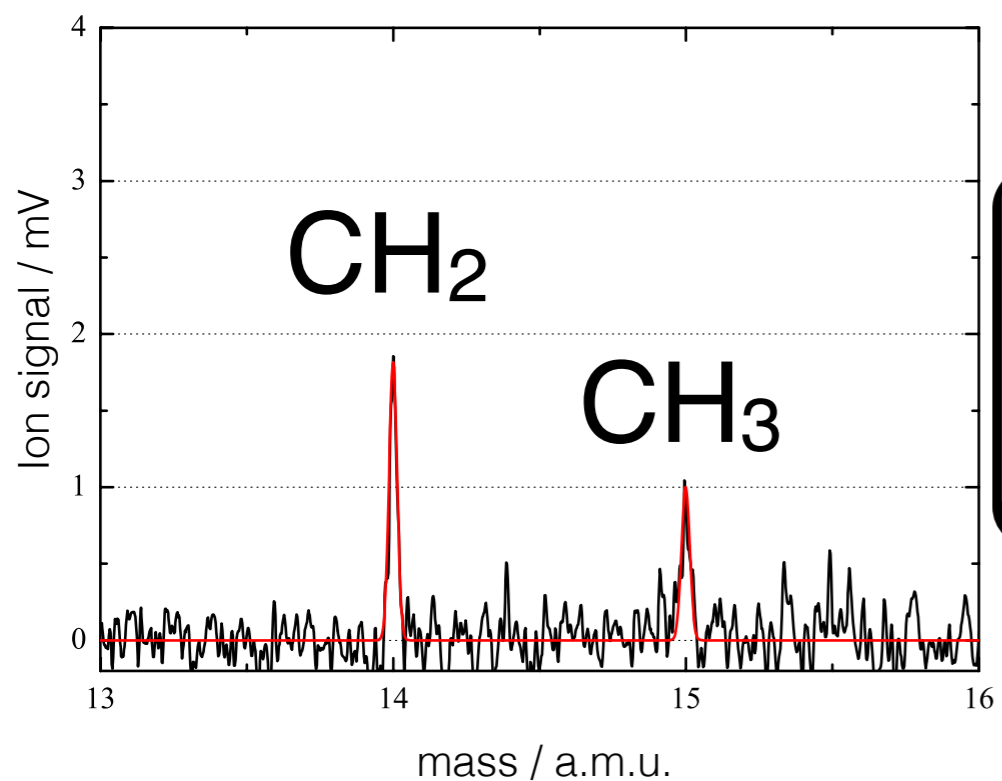
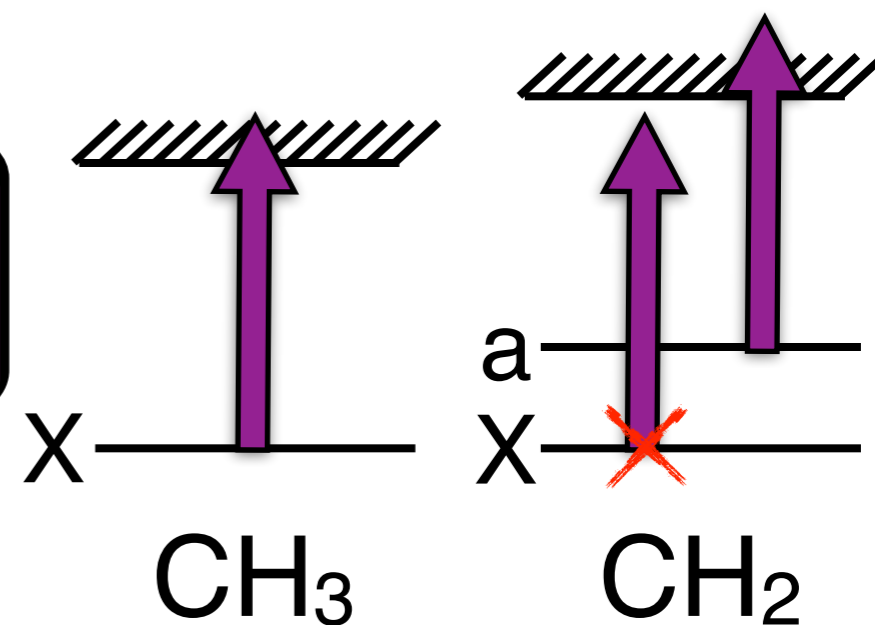


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Ionization at 118.2 nm (10.49 eV)	✓	✓	✓	✗	✓

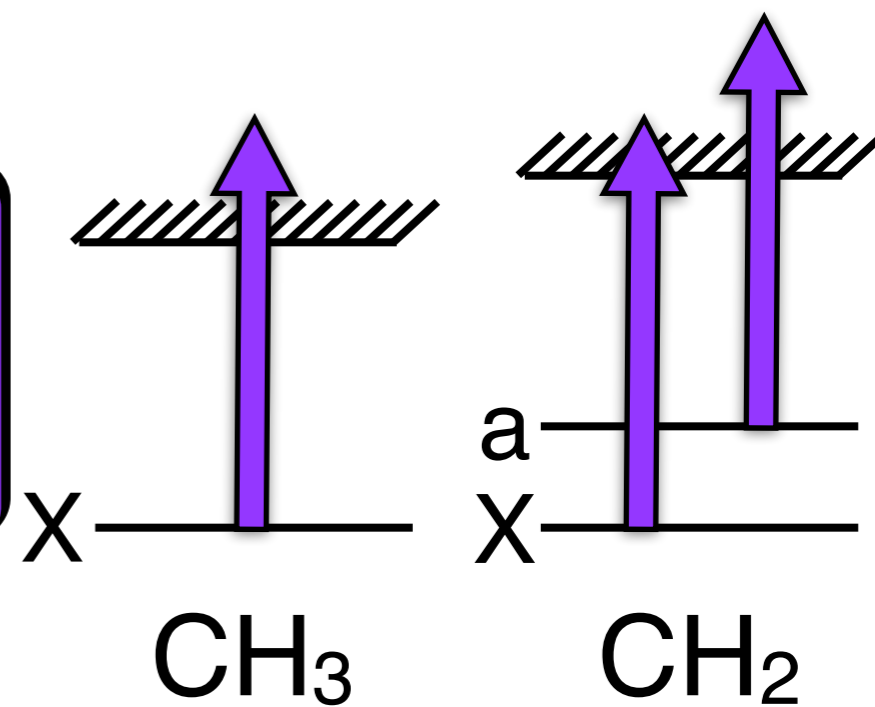
Photolysis at 121.6 nm

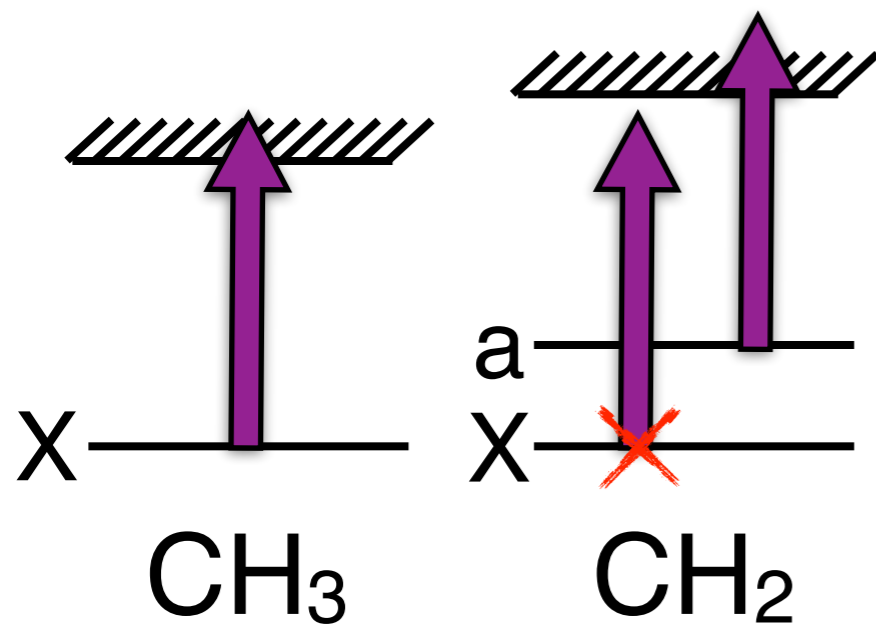


Ionization  
at 121.6 nm  
CH<sub>2</sub>(a) & CH<sub>3</sub>(X)



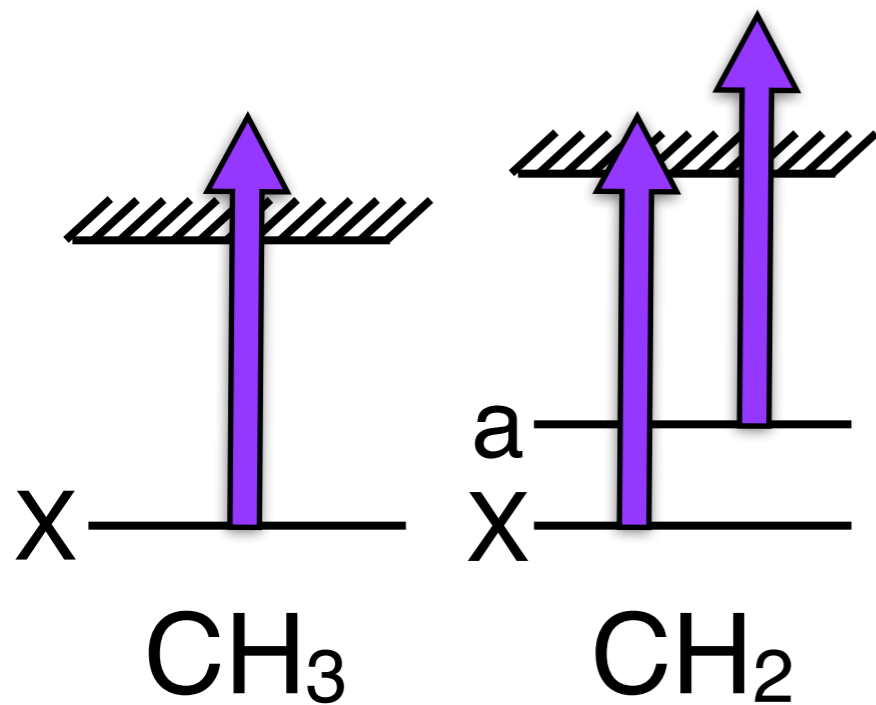
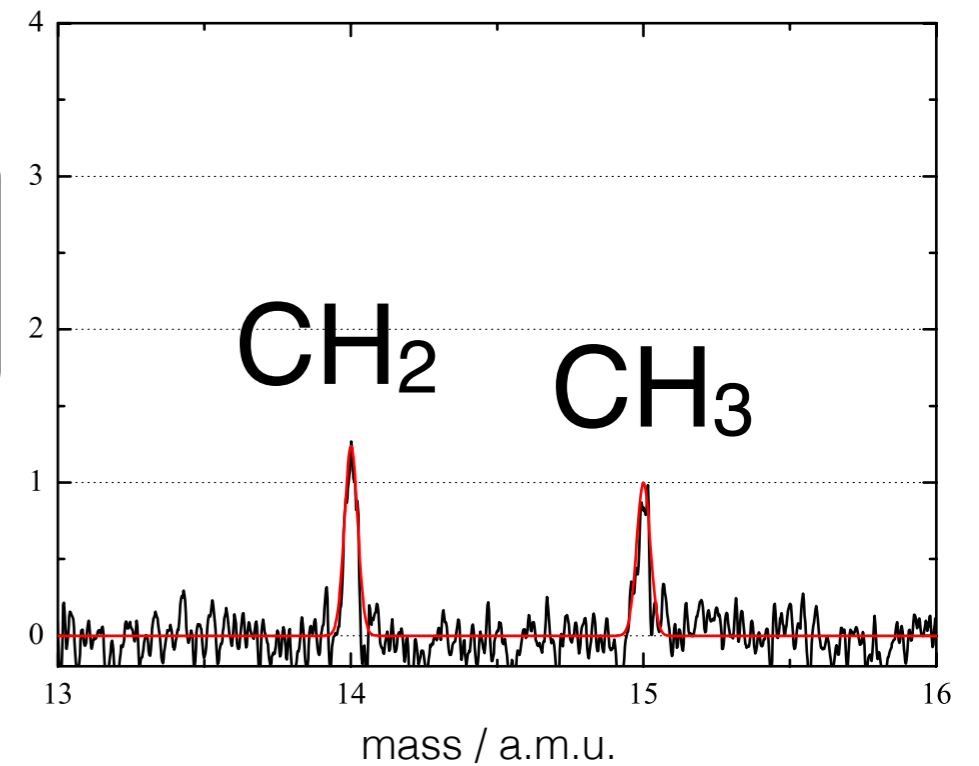
Ionization  
at 118.2 nm  
CH<sub>2</sub>(a), CH<sub>3</sub>(X) &  
CH<sub>2</sub>(X)



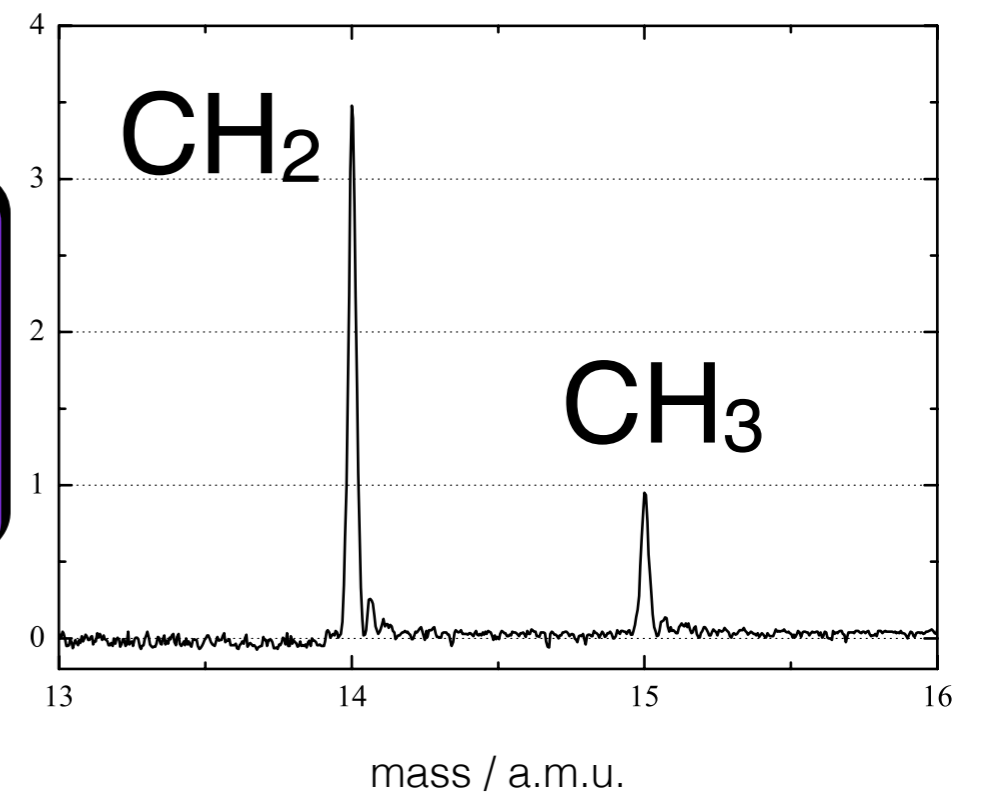


Ionization  
at 121.6 nm  
 $\text{CH}_2(\text{a})$  &  $\text{CH}_3(\text{X})$

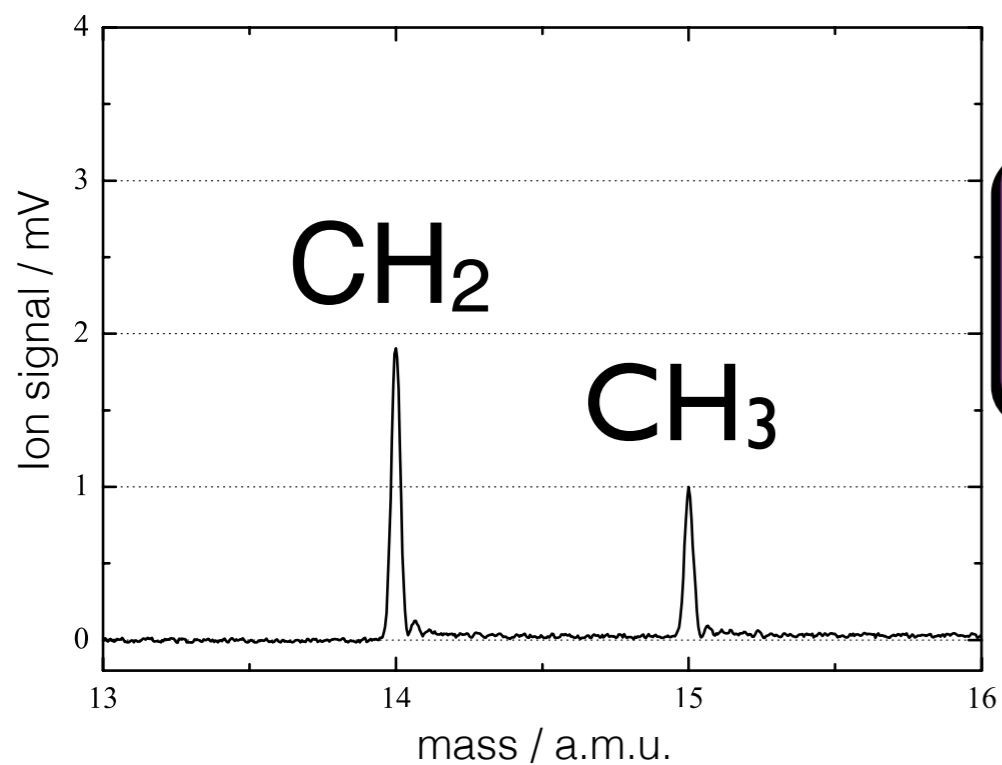
Photolysis at 118.2 nm



Ionization  
at 118.2 nm  
 $\text{CH}_2(\text{a})$ ,  $\text{CH}_3(\text{X})$  &  
 $\text{CH}_2(\text{X})$

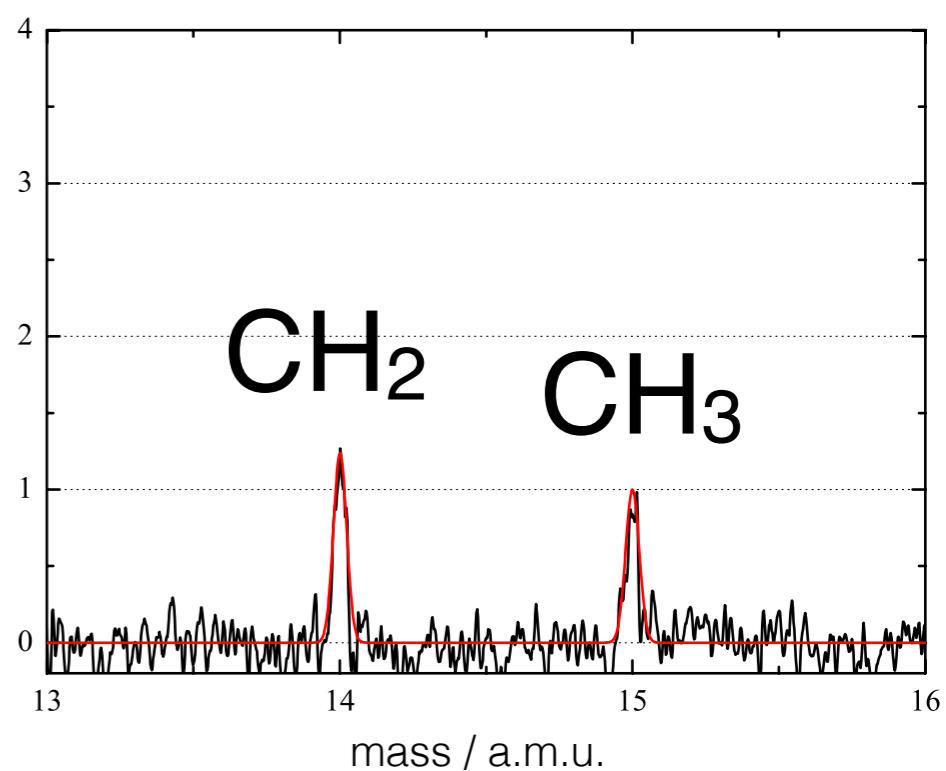


Photolysis at 121.6 nm

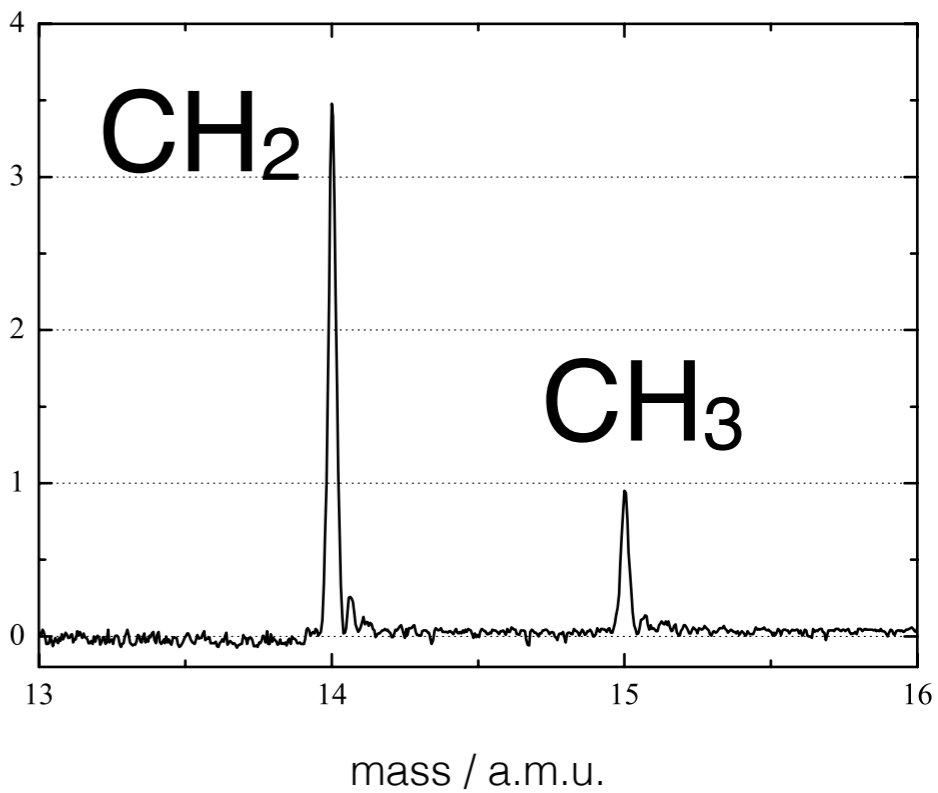
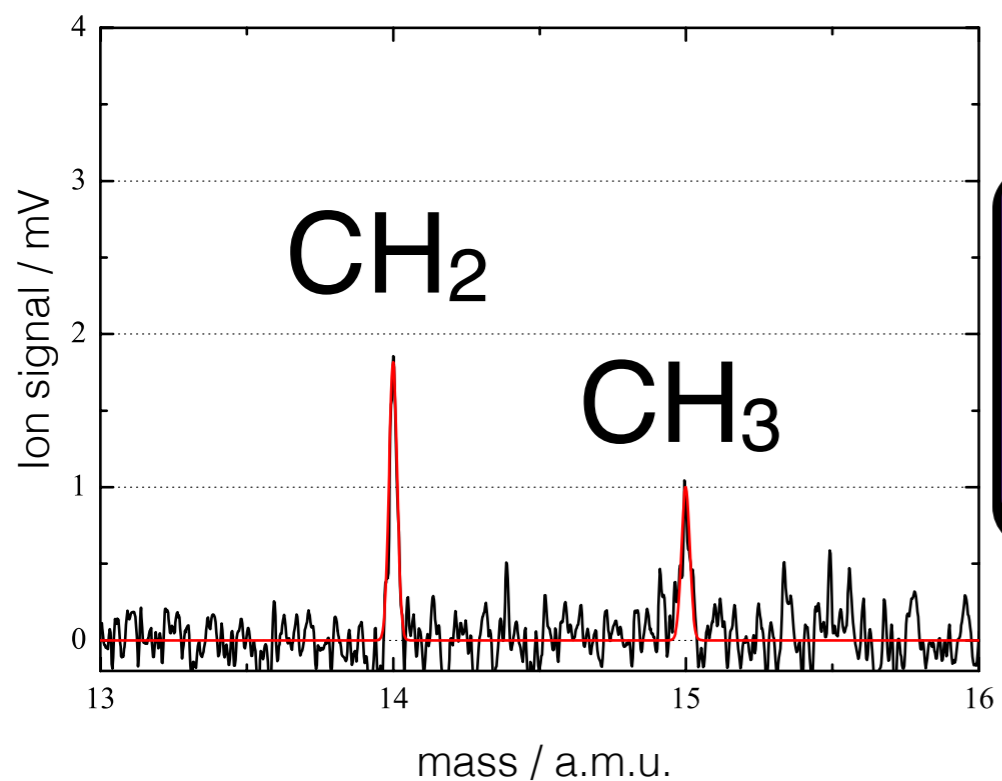


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at 121.6 nm  
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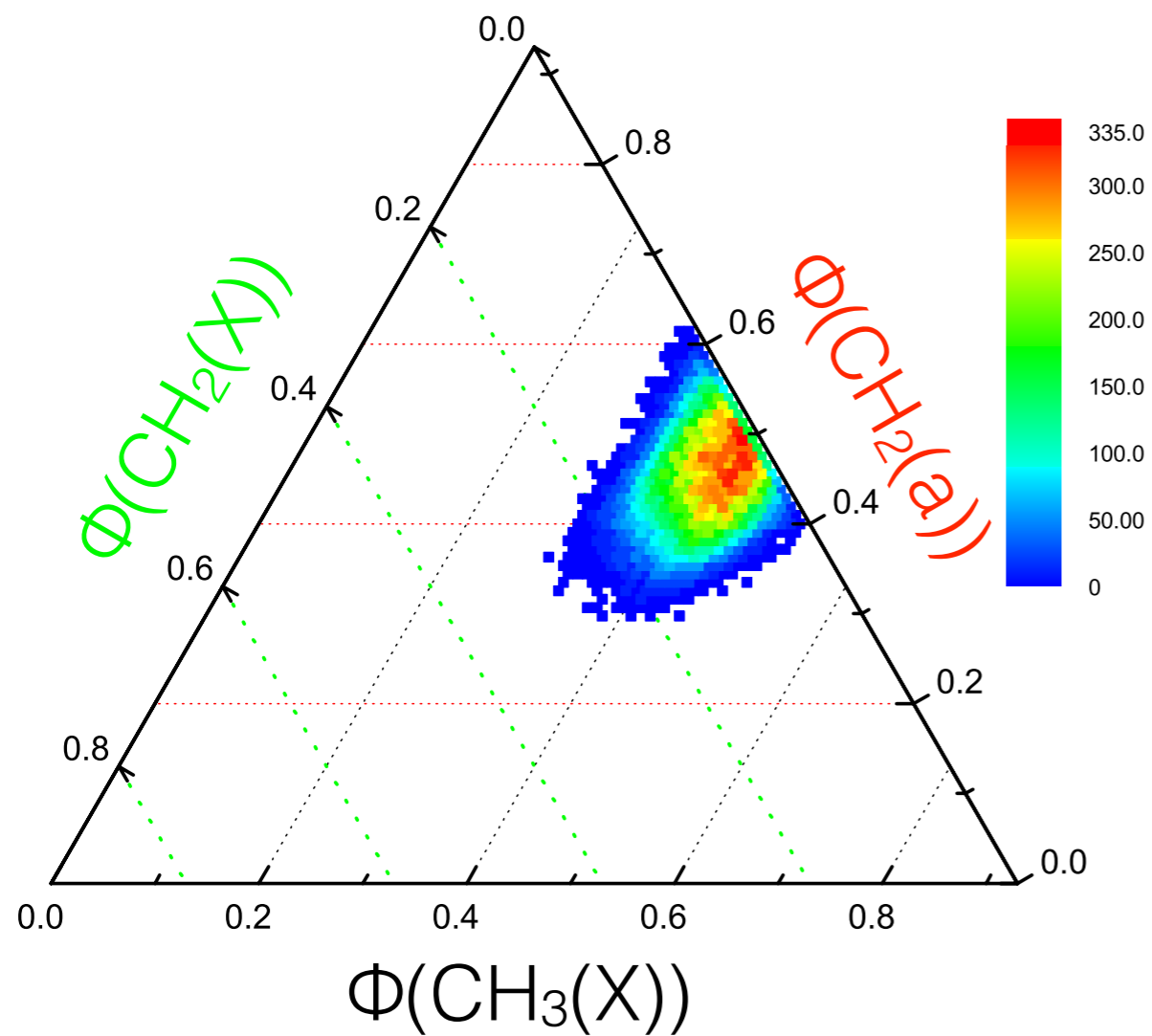


Ionization  
at 118.2 nm  
CH<sub>2</sub>(a), CH<sub>3</sub>(X) &  
CH<sub>2</sub>(X)



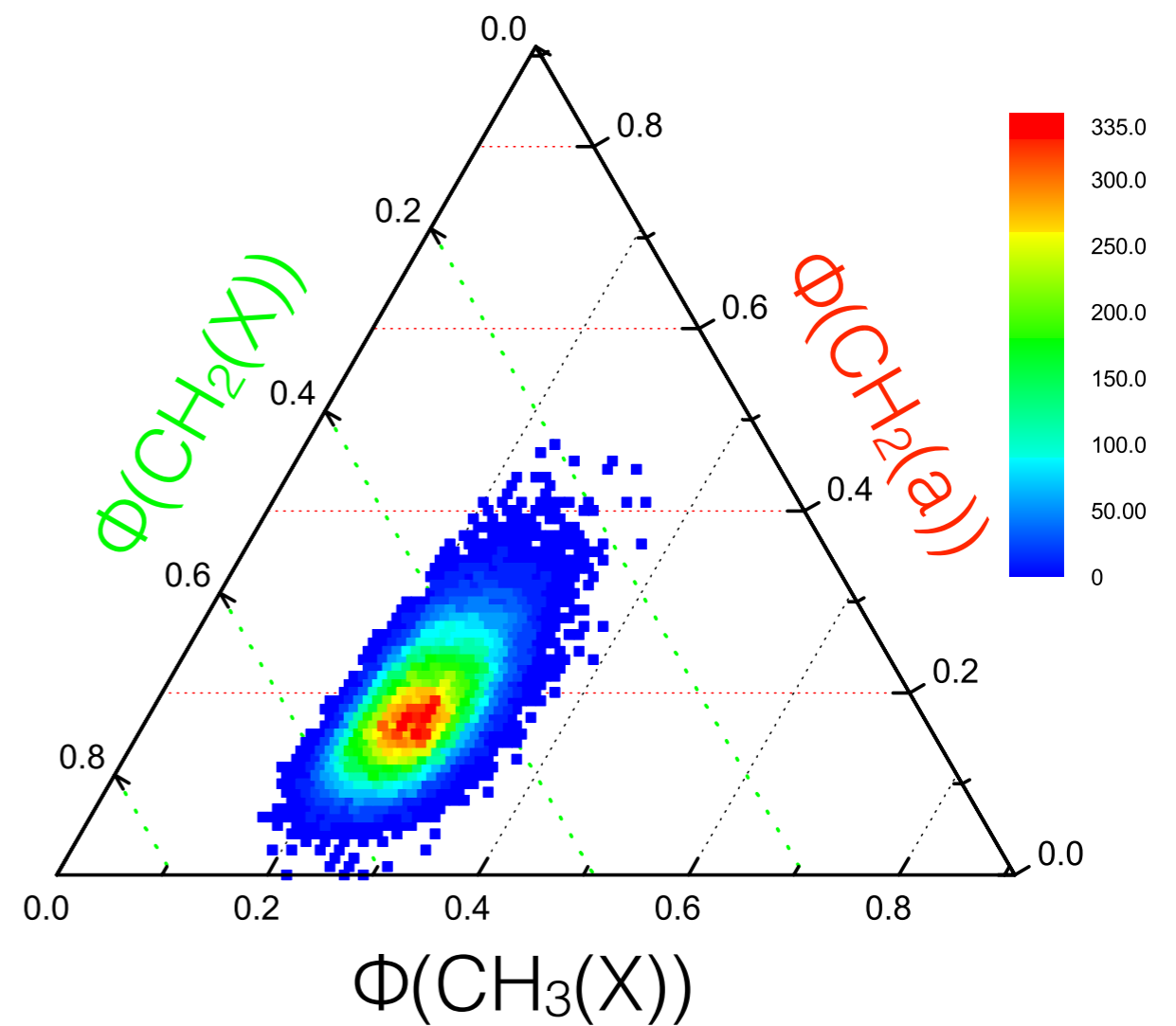
# Results: branching ratios

Photolysis at 121.6 nm



$\Phi(\text{CH})=0.07$

Photolysis at 118.2 nm



$\Phi(\text{CH})=0.09$

■  $\text{CH}_3(\text{X})$  ■  $\text{CH}_2(\text{X})$  ■  $\text{CH}_2(\text{a})$

# Comparison with previous studies

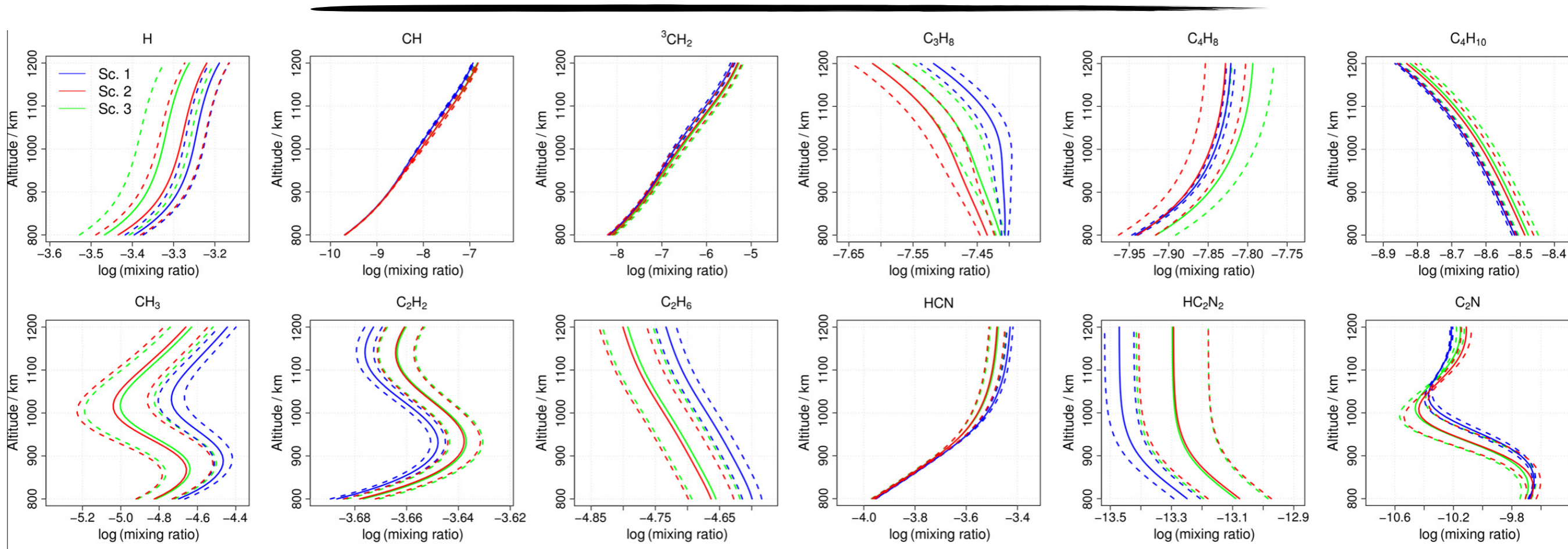
Dissociative channel	Mordaunt et al. S1 & S2 1993		Heck et al. (1995)	Brownsword et al. (1997)	Wang et al. (2000)	Park et al. (2008)	This work
CH <sub>3</sub> (X)+H	0.51	0.49	0.66	0.38	0.291	0.31	0.42
CH <sub>2</sub> (a)+H <sub>2</sub>	0.24	0	0.22	0.52	0.584	0.63	0.48
CH <sub>2</sub> (b)+H <sub>2</sub>	/	/	/	/	/	/	≈0
CH <sub>2</sub> (X)+2H	0.05	≈0	≈0	0.01	/	/	0.03
CH <sub>2</sub> (a)+2H	0.2	≈0	≈0		0.055	/	≈0
CH(X)+H+H <sub>2</sub>	0	0.51	0.11	0.08	0.07	0.059	0.07
C( <sup>1</sup> D)+2H <sub>2</sub>	0	0	0	0	0	0.0004	≈0
Φ(H)	1	1	0.77	0.47	0.47	0.31	0.55
Φ(H <sub>2</sub> )	0.24	0.51	0.33	0.6	0.654	0.69	0.55

# Conclusions

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- Pump-probe experiment with 2 VUV beams to study methane photolysis (121.6 nm & 118.2 nm)
- New set of branching ratios at 121.6 nm and first set of BR at 118.2 nm.
  - ➔ Strong wavelength-dependence of the BR.
- Ready to be implemented in KIDA database !!
- New theoretical studies should be stimulated by these experimental results!

# Test with a photochemical model of Titan



B. Gans, Z. Peng, N. Carrasco, D. Gauyacq, S. Lebonnois, P. Pernot, *Icarus* **223** (2013) 330–343

⇒ wavelength dependence of branching ratios should be taken into account.



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# Perspectives

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- Implementation of CH<sub>4</sub> photolysis BR in KIDA database
- Expertise in BR measurements and radical cross sections to be used to study other astrophysically-relevant species:
  - UV and VUV photolysis of NH<sub>3</sub>
  - UV and VUV photolysis of cyanopolyynes (HC<sub>3</sub>N,....)
  - ...

Thank you for your attention!

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